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TRANSPORT PROPERTIES IN THE ATMOSPHERE OF JUPITER

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#### INTRODUCTION

The work discussed in this report is concerned with the calculation of transport properties near the surface of a probe entering the atmosphere of Jupiter. The discussion of this work is divided into the following categories; (1) transport properties in the pure Jovian atmosphere, (2) transport properties for collisions between monatomic carbon atoms, including the effect of excited electronic states, (3) transport properties at the boundaries for mixing of the pure Jovian atmosphere and the "atmosphere" due to the injection of gaseous ablation products, and (4) transport properties for interactions involving some of the molecular ablation products.

The transport properties are calculated using the kinetic theory of gases.

This theory is well developed for elastic collisions involving neutral atoms and/or small polyatomic (usually diatomic) species. The theory is in reasonably good shape for collisions involving ions. The calculation of the contribution of inelastic collision effects to the transport properties is still quite difficult.

The determination of the interaction potential between the interacting atoms/ions/molecules is usually the primary problem in the calculation of the transport properties. Transport collision integrals have been calculated for only a limited set of empirical and semiempirical interaction potentials. Since the accuracy of the fit of these empirical potentials to the "true" potential usually determines the accuracy of the calculation of the transport properties, a discussion of the various interaction potentials used in these calculations will be emphasized in this report.

### II. TRANSPORT PROPERTIES IN THE PURE JOVIAN ATMOSPHERE

The nominal chemical composition of the Jovian atmosphere is taken to be  $X_{H_2} = 0.89$  and  $X_{H_e} = 0.11$ , where X denotes mole fraction. If it is assumed that the atmosphere is at chemical equilibrium, the mole fractions of various species as

a function of temperature given in Table 1 are obtained.<sup>5</sup> The table lists temperatures to 25,000°K since the calculations of Moss, et al.,<sup>6,7</sup> indicate that high temperatures are attained near a probe upon entry into the Jovian atmosphere.

# A. The Interaction Potentials

Transport properties for the species listed in Table 1 have been reported. 8

The discussion given in that paper will not be repeated in this report. However, it is useful to review the interaction potentials used in the calculations since the potentials are the single most important "ingredient" in the calculation.

Many of the species in the Jovian atmosphere interact according to a potential which is repulsive at short range but possesses an attractive potential well at intermediate separations. Such interactions have been represented by one of the empirical (or semiempirical) potentials given below; (a) the attractive inverse power (AIP) potential,

$$V(r) = -\frac{A}{n^{B}} \tag{1}$$

where V is the potential energy, r is the internuclear separation, and A and B are adjustable parameters, (b) the exponential-six (ES) potential,

$$V(r) = \frac{\varepsilon}{1 - \frac{6}{\alpha}} \left[ \frac{6}{\alpha} e^{\alpha (1 - r/r_e)} - \left( \frac{r_e}{r} \right)^6 \right]$$
 (2)

where  $\varepsilon$  is the depth of the poential well,  $r_e$  is the value of r when  $V = -\varepsilon$ , and  $\alpha$  is an adjustable parameter, or (c) the Morse potential (MP),

$$V(r) = \varepsilon \left[ e^{-2 \frac{c}{\sigma} (r - r_e)} - 2e^{-\frac{c}{\sigma} (r - r_e)} \right]$$
 (3)

where c is an adjustable constant and

Other species in the Jovian atmosphere interact according to a potential which is repulsive at all separations. Such interactions have been represented by the exponential repulsive (ER) potential;

$$V(r) = Fe^{-Dr} \tag{4}$$

where D and F are adjustable constants.

Transport collision integrals have been calculated and tabulated for each of the potentials given above; AIP-reference 9, ES-reference 10, MP-reference 11, and ER-reference 12. Results for the interactions of interest in the Jovian atmosphere are summarized in Table 2. The third column lists the papers in which the "ab initio" calculations of the interactions listed in the first column are discussed and the fourth column lists the papers in which a "best fit" of the empirical potentials to the ab initio results is discussed.

In addition, the screened Coulomb potential is used for the ion-ion, ion-electron, and electron-electron interactions. For ions with unit electrical charge, the potential has the form  $^{23}$ 

$$V(r) = \pm \frac{e^2}{r} e^{-r/\lambda} d$$
 (5)

where e is the electrical charge and  $\boldsymbol{\lambda}_d$  is the Debye shielding distance, given by

$$\lambda_{d} = \left(\frac{kT}{4\pi e^{2}n_{e}}\right)^{1/2}$$

and  $n_e$  is the electron density. The transport collision integrals have been tabulated for this potential.  $^{23,24}$ 

Also, the H-H<sup>+</sup> and He-He<sup>+</sup> diffusion collision integrals were obtained by considering resonant charge exchange, the collision integrals for the H-e interaction were obtained from data on low energy elastic scattering cross sections, and the He-e collision integrals were obtained from data on the diffusion cross section.<sup>8</sup>

The resulting transport collision integrals for the interactions occuring in the pure Jovian atmosphere are shown in Tables 3 to 17. These collision integrals have been used to calculate the transport properties of each component and of the gas mixture.

### B. Errors in the Interaction Putentials

As indicated previously, the primary source of error is in the fitting of the empirical potentials for which the transport collision integrals are tabulated to the ab initio potentials, for which the collision integrals have not been calculated.

An example of the fit is shown in Table 18 for the He-H $^+$  interaction. The results in the second column are a representation of the accurate quantum mechanical calculations of Evett $^{18}$  near  $r_e$  and the somewhat less accurate estimates of Mason and Vanderslice $^{22}$  at large values of r. The results in the third column represent the "best fit" of the Morse potential to the ab initio results. Clearly the fit is relatively poor at small and large values of r but quite good near  $r = r_e$ . The reason for this is that the attractive region of the potential dominates the scattering process when r

$$T^* = \frac{kT}{\epsilon} < \sim 5 \text{ or } 6$$

For the He-H<sup>+</sup> interaction,  $T^* = 1.142$  when T = 25,000°K. Thus the Morse parameters have been chosen so that the best fit to the ab initio results is at  $r = r_o$ .

The results in Table 18 are quite typical although, for some interactions, the curve fit is much poorer. Recent work concerned with improving the fit of the empirical potentials to the ab initio results will be discussed later.

#### C. Errors in the Ion-Ion Interactions

The second major source of error in determining the transport properties in the pure Jovian atmosphere is a consequence of the approximations used in the calculation of the ion-ion collision integrals. The Chapman-Enskog method was used to calculate these collision integrals, using the screened Coulomb potential and assuming static screening by electrons only. Also, only the lowest order approximation was used in calculating the collision integrals.

The lowest order approximation may be satisfactory for a highly ionized gas, in which case dynamic shielding results are very similar to the results obtained for static shielding by electrons only. 25 However, the third 25,26 or higher 27 order terms in the Chapman-Enskog series should probably be included in the calculation in order to get good convergence to the "true" results for the transport properties. The convergence is particularly slow for transport properties that depend primarily on the electron distribution (such as the electron-electron and electron-ion binary diffusion coefficients, the thermal conductivity at high ionization, and the electrical conductivity) and is "rapid" for transport properties that depend primarily

on the heavy particle (atoms and ions) distributions (such as the viscosity, the thermal conductivity at low ionization, and the heavy particle binary diffusion coefficients).

At low ionization (less than  $\sim 10\%$ ), the convergence problems are even greater.  $^{25,28}$  Under these conditions, the ionized gas approaches a Lorentzian mixture (i.e. a few light particles and many heavy, stationary particles). It has been shown that the calculation of the transport properties by considering a perturbation on the Lorentzian distribution is more efficient than the Chapman-Enskog approach (i.e. it converges rapidly).

The corrections at low and high ionization discussed above should be made for the ion and electron transport collision integrals. Also, the electric and magnetic fields in the Jovian atmosphere due to the presence of ions and electrons (and other causes) have an effect on the transport properties.

# III. TRANSPORT PROPERTIES OF MONATOMIC CARBON

Entry probe heat shields are usually made of carbonaceous materials such as carbon-phenolic ablator: typically 92% carbon, 6% oxygen, and 2% hydrogen by mass. Ablative injection of the carbon-phenolic material into the shock layer is an important mechanism for reducing the intense radiative heating encountered during entry into the atmospheres of the outer planets. The species C,  $C_2$ , and  $C_3$  are important ablative species, especially at low temperatures and near the entry probe. This is shown in Table 19. Note that  $C_3$  is the predominant specie near the surface of the probe. Thus it is particularly important that accurate estimates of the transport collision integrals be obtained for interactions involving  $C_3$ .

Transport collision integrals have been calculated for the interaction of two ground state carbon atoms. These calculations will be described in some detail since the results will be used to obtain collision integrals for the C-C<sub>2</sub>, C<sub>2</sub>-C<sub>2</sub>, C-C<sub>3</sub>, C<sub>2</sub>-C<sub>3</sub>, and C<sub>3</sub>-C<sub>3</sub> interactions. Thus it is important to obtain the best possible estimate of the C-C transport collision integrals (which means obtaining

the best possible estimate of the C-C interaction potentials).

A. Ground State Interaction Potentials

When two ground state ( $^3$ P) carbon atoms interact, they can follow  $^{32}$  any of 18 interaction curves, corresponding to 18 molecular states of  $^2$ C. Accurate interaction potentials are needed for each of these 18 states. Thirteen of the states possess an attractive potential well (bound states) and the spectroscopic constants (i.e. the dissociation energy,  $\epsilon$ , the internuclear separation at the minimum in the potential well,  $r_e$ , the fundamental vibrational frequency,  $\omega_e$ , the rotational constant,  $B_e$ , the anharmonicity constant,  $\omega_e x_e$ , and the rotation-vibration coupling constant,  $\alpha_e$ ) have been either measured experimentally  $^{33}$  or accurately estimated.  $^{34,35}$  Five of the states are repulsive states.

These 18 states are listed in Table 20. The last five states are the repulsive states.

The interaction potentials for the 13 bound states can be represented by the Hulburt-Hirschfelder (HH) potential,  $^{36,37}$  given by

$$V(r) = \varepsilon[(1-e^{-x})^2 + Cx^3e^{-2x} (1+bx)]$$
 (6)

where

$$x = \frac{\omega_{e}}{2r_{e}\sqrt{B_{e}\varepsilon}} \quad (r-r_{e}) \qquad C = 1 + a_{1}\sqrt{\frac{\varepsilon}{a_{0}}}$$

$$b = 2 - \frac{7/12 - \varepsilon a_{2}/a_{0}}{C} \qquad a_{0} = \frac{\omega_{e}^{2}}{4B_{e}}$$

$$a_{1} = -1 - \frac{\alpha_{e}\omega_{e}}{6B_{e}^{2}} \qquad a_{2} = \frac{5}{4}a_{1}^{2} - \frac{2\omega_{e}\chi_{e}}{3B_{e}}$$

This empirical potential uses the six spectroscopic constants as parameters and is nearly as accurate a representation of the "true" potential energy curve as is available  $^{36,38,39}$  (alternatives will be discussed later). It also has the intellectually satisfying feature that all parameters are fixed by experiment; i.e. it has no adjustable parameters.

The Hulburt-Hirschfelder potential energy curves were obtained for each of the first 13 states listed in Table 20. Unfortunately, until recently (recent o velopments will be discussed later), transport collision integrals had not been calculated for this potential. This the Hulburt-Hirschfelder curves were best fit with the Morse potential, as described previously. The best fit parameters are shown in the third column of Table 20.

Theoretical results  $^{34}$  for three of the repulsive states were best fit with the exponential repulsive potential. The best fit parameters are also shown in the third column of Table 20. The  $^3\Sigma^+_{u2}$  and  $^5\Sigma^+_{g2}$  states were assumed to be degenerate with the  $^3\Sigma^+_u$  and  $^5\Sigma^+_g$  states, respectively, although, as will be discussed later, this assumption can be avoided.

Transport collision integrals for each of the Morse and exponential repulsive curves were calculated and then averaged according to their degeneracies.  $^{16}$  The results are shown in Table 21. These results were used  $^{31}$  to calculate the transport properties in a gas of  $^{3}$ P carbon atoms.

1. Use of the Hulburt-Hirschfelder Potential

It is important to emphasize again that an accurate  $C(^3P)-C(^3P)$  interaction potential is highly desirable since this potential will be used as the basis for constructing interaction potentials for interactions involving C,  $C_2$ , and  $C_3$ .

The primary source of error for the  $C(^3P)$ - $C(^3P)$  interaction potentials is due to errors in the curve fit of the Morse potential to the accurate Hulburt-Hirschfelder potential. An example of the "goodness" of the curve fit is shown in Table 22, for the  $^1\Sigma_{\bf g}^+$  state. As before, the fit has been optimized near  ${\bf r}_{\bf e}$  but, at large and small values of  ${\bf r}$ , the two potentials are quite different. This leads to errors in the calculation of the transport collision integrals.

Until recently, tabulations of collision integrals for the Hulburt-Hirschfelder potential were not available. However, a previously developed program for calculating collision integrals has now been modified and adapted for the Hulburt-

Hirschfelder potential. Indeed, it will even calculate collision integrals for states with a "wiggle" at large values of r such as the  ${}^1\Pi_g$  state of  $C_2$  shown in Table 23. Such wiggles may be real and not simply an artifact of the potential. <sup>36</sup> They seem to occur often for the states of  $C_2$ . <sup>34</sup> One possible cause of the wiggles is rotational instability <sup>32,41</sup> in  $C_2$ .

Thus transport collision integrals can now be calculated for the Hulburt-Hirschfelder potential. Results for the states of  $\mathbf{C}_2$  are now being calculated. Results for the appropriate interactions in the pure Jovian atmosphere will also be calculated.

Some results are available. Table 24 lists the transport collision integrals for the  $^{1}\Sigma_{g}^{+}$  state of  $C_{2}$  for the Hulburt-Hirschfelder and the Morse potentials. The differences are quite substantial. These differences are directly reflected in the calculation of the transport properties of  $^{3}P$  carbon atoms. They will also be reflected (although not directly) in the calculation of transport properties in a gas mixture containing  $^{3}P$  carbon atoms.

The value of  $\sigma$  for the Hulburt-Hirschfelder potential was taken to be the value of r (other than  $r=\infty$ ) for which V(r)=0. It was calculated by an iterative method, using the relation

$$2e^{A(t-1)} = B(t-1)^{3}[1 + G(t-1)] + 1$$

where

$$t = \frac{\sigma}{r_e}$$
  $A = \frac{\omega_e}{2\sqrt{B_e \varepsilon}}$   $B = cA^3$   $G = bA$ 

For the Hulburt-Hirschfelder potential,  $\sigma$  = 0.9494Å while  $\sigma$  = 0.9724Å for the Morse potential. Thus most of the difference in the results for the two potentials is in the calculation of  $\Omega^{(1,1)*}$  and  $\Omega^{(2,2)*}$ .

How accurate is the Hulburt-Hirschfelder potential? While other potentials, with adjustable parameters, may give a better fit to the "true" potential for specific chemical species, probably no other empirical or semiempirical potential gives a better fit to true potentials for a wider range of chemical species. 38

It should be pointed out that, while the Hulburt-Hirschfelder potential is based on experimental information, it is also model dependent since the spectroscopic constants are calculated by assuming that molecules are anharmonic vibrators and non-rigid rotors. However, a method is available for determining potential energy curves which makes use of the experimental energy levels directly (i.e. it is not model dependent) and also does not depend on assuming a functional form for the potential; i.e. the results are obtained in the form of a table of V(r) versus r. This method was developed by Rydberg,  $^{42}$  Klein,  $^{43}$  and Rees  $^{44}$  and is called the RKR method. It gives results that agree with the results obtained by the Dunham method near  $r_e$ .  $^{46,47}$  The RKR method is the most accurate method for determining the interaction potential energy between atoms that interact according to a long range attractive potential.

Table 25 gives a comparison of the RKR results  $^{48}$  with the Hulburt-Hirschfelder results for the  $^1\Sigma_g^+$ ,  $^3\Pi_u$ , and  $^1\Pi_g$  states of  $C_2$ . The agreement tends to be very good near  $r_e$  but differences can be substantial at large values of r. These results, and others, indicate that the Hulburt-Hirschfelder potential is quite accurate, especially near  $r_e$ , the region that usually makes the predominant contribution to the scattering.

It should be pointed out that the computer program that has been developed can take V(r) versus r "data", fit it with a polynomial, and calculate the transport collision integrals for the resulting polynomial curve fit; i.e. an assumed functional form for the potential is not required. Thus transport collision integrals can be obtained directly from the RKR results, if necessary.

### 2. The Perfect Pairing Nathod

It was previously assumed  $^{31}$  that the  $^{3}\Sigma_{u2}^{+}$  and  $^{5}\Sigma_{g2}^{+}$  states are degenerate with the  $^{3}\Sigma_{u}^{+}$  and  $^{5}\Sigma_{g}^{+}$  states, respectively, since good quantum mechanical calculations of the interaction energy are not available for these states. However, an approximate method of estimating the potential energy for these states based on valence bond theory,  $^{49}$  called the perfect pairing method, is available. This method will be described briefly.

If it is assumed that eight  $\epsilon$ lectrons in  $C_2$  fill the four lowest molecular orbitals, i.e.

$$(\sigma_{q}1s)^{2}(\sigma_{u}^{*}1s)^{2}(\sigma_{q}2s)^{2}(\sigma_{u}^{*}2s)^{2}$$

then, among the possible arrangements for the four remaining electrons, are the arrangements corresponding to the  $^3\Sigma_{u2}^+$  and  $^5\Sigma_{g2}^+$  states $^{50,51}$  shown in Table 26, along with arrangements for the  $^3\eta_u$  and  $^3\Sigma_g^-$  states. The direction of each arrow corresponds to the "direction of the spin".

The energy relationships are simple. They are based on a valence bond treatment that is correct at large values of r. The Coulomb integral is assumed to make no contribution to the energy and there is a contribution to the energy of  $\frac{1}{2}$  J (J is the exchange integral) from each electron in a bonding molecular orbital and contribution of  $-\frac{3}{2}$  J from each electron in an antibonding orbital. 51 Also

$$J_{xx} = J_{yy}$$

Thus

$$V(^{3}\Sigma_{u2}^{+}) = -J_{zz} + J_{xx} \qquad V(^{5}\Sigma_{q2}^{+}) = -2J_{xx}$$

The integrals  $J_{xx}$  and  $J_{zz}$  must be evaluated using information about two states for which potential energy curves are known. The  $^3\mathrm{H}_u$  and  $^3\mathrm{E}_g^-$  states have been used for this purpose since accurate RKR results are available for these states  $^{48}$  and the RKR results are accurately reproduced by the Hulburt-Hirschfelder potential. For these states

$$V(^{3}\Pi_{u}) = \frac{1}{2}J_{zz} + \frac{3}{2}J_{xx}$$
  $V(^{3}\Sigma_{g}^{-}) = J_{zz} + J_{xx}$ 

or

$$J_{xx} = \frac{2V(^{3}\Pi_{u}) - V(^{3}\Sigma_{g}^{-})}{2} \qquad \qquad J_{zz} = \frac{3}{2}V(^{3}\Sigma_{g}^{-}) - V(^{3}\Pi_{u})$$

Using the known values of  $V(^3II_u)$  and  $V(^3\Sigma_g^-)$ , the quantities  $J_{\chi\chi}$  and  $J_{\chi\chi}$ , and thus  $V(^3\Sigma_{u2}^+)$  and  $V(^5\Sigma_{g2}^+)$ , can be estimated. The results are snown in Table 27. These results can be curve fit with one of the empirical repulsive potentials and the transport collision integrals can then be determined.

When the calculations discussed in sections III-A-1 and III-A-2 are completed, a substantially improved estimate of the transport properties of ground state carbon atoms will be available.

# B. Excited State Interaction Potentials

The temperatures attained during entry of a probe into the Jovian atmosphere are so high that a significant fraction of the carbon atoms are in excited electronic states. This is shown in Table 28, obtained by including all of the electronic states listed in the JANAF Tables in the calculation and assuming that there is an equilibrium distribution of atoms among the states. Clearly, at temperatures above 6000°K, excited electronic states are significantly populated.

In previous reports, <sup>52,53</sup> possible models for calculating the contribution to the transport properties from species in excited electronic states were discussed. Since the results obtained using these models are less reliable than the results to be presented below, these models will not be discussed in this report. However, the model calculations do confirm the assumption <sup>54</sup> that the contribution from low lying excited electronic states to the transport properties is nearly the same as the contribution from the ground state and that the contribution from highly excited electronic states to the transport properties is negligible.

Now consider the interaction between a ground state ( $^3$ P) carbon atom and a carbon atom in the first excited ( $^1$ D) electronic state. The molecular states of  $C_2$  that dissociate into a  $^3$ P carbon atom and a  $^1$ D carbon atom are shown  $^{32,34}$  in the first column of Table 29. Experimental values of the spectroscopic constants are available  $^{32}$  for the  $^3$ II $_{g3}$  state. For the other bound states (except the  $^3$ II $_{u2}$  state), theoretical results  $^{34,35}$  have been correlated in order to obtain reasonable estimates of the spectroscopic constants.  $^{35}$  The results are shown in Table 29. The  $^3$  $\varphi_u$ ,  $^3$  $\Sigma_g^+$ , and  $^3$  $\Sigma_u^-$  repulsive states were investigated theoretically by Fougere and Nesbet.  $^{34}$ However, the level of refinement of their calculations for these states is relatively crude.

For the  $^3\pi_{g2}$  and  $^3\pi_{u2}$  bound states, Morse parameters were obtained by making a best fit of the Morse potential to the theoretical results of Fougere and Nesbet.  $^{34}$  For the other bound states, the Morse parameters were obtained by making a best fit of the Morse potential to the Hulburt-Hirschfelder potential.

In principle, collision integrals for the repulsive states should be obtained by using an empirical repulsive potential. However, while at the level of calculation III by Fougere and Nesbet,  $^{34}$  the  $^{3}\Phi_{u}$ ,  $^{3}\Delta_{g}$ , and  $^{3}\Sigma_{u}^{-}$  states are repulsive, at the level of their least accurate calculation (called calculation I--the only level at which results are available for these states), these three states exhibit a shallow attractive minimum in the potential. Thus Fougere and Nesbet's results for these states  $^{34}$  have been best fit with the Morse potential, using the "spectroscopic constants" listed in Table 30. Clearly this is a rather crude approach.

The parameters obtained for the empirical potentials are given in Table 30. Notice that the last eight states listed in Table 30 have not been included in the calculation. This is a serious source of error. However, a crude estimate of the potential energy curve for each of these states can be obtained by using the perfect pairing method, discussed previously.

Transport collision integrals can be obtained for each of the first ten states listed in Table 30 and the integrals can then be averaged according to their degeneracies. <sup>16</sup> The results are given in Table 31. In addition to the errors due to ignoring the contribution to the collision integrals from eight states, the other sources of error discussed in connection with the  $C(^3P)-C(^3P)$  calculation are sources of error for this calculation.

Now consider the interaction between two  $^1D$  carbon atoms. The molecular states of  $C_2$  that dissociate into two  $^1D$  carbon atoms are shown  $^{34}$  in the first column of Table 32. Experimental spectroscopic constants are not available for any of these states. The estimated spectroscopic constants  $^{34}$  for the bound states are also shown in Table 32. At the level of calculation III by Fougere and Nesbet,  $^{34}$  the  $^1\Gamma_G$  and

 $\Phi_{\mathbf{u}}$  states are repulsive, but they exhibit a shallow attractive minimum at the level of calculation I.

The first seven states listed in Table 32 have been best fit with the Morse potential. The resulting parameters are given in Table 32. The last eight states listed in Table 32 have been ignored in the calculation. This is, of course, a serious source of error but the perfect pairing method can be used to obtain estimates of the potential energy curves for these repulsive states.

Transport collision integrals for the first seven states listed in Table 32 have been obtained and averaged according to their degeneracies. <sup>16</sup> The results are given in Table 33. The sources of error discussed previously also apply to these results.

It is clear that experimental and/or theoretical information for the excited electronic states of carbon is rather limited. Thus the results in Tables 31 and 33 must be considered to be relatively crude first order estimates. However, the results are almost certainly accurate to within less than a factor of 2 and the inclusion of perfect pairing results for the omitted states will probably not change the ransport collision integrals significantly.

The results in Table 33 can be used to calculate the transport properties in a gas of  $^1D$  carbon atoms. It is of greater interest, however, to use the results in Tables 21, 31 and 33 to calculate  $^{1,8}$  the transport properties in a mixture of  $^3P$  and  $^1D$  carbon atoms. Some results for the translational contribution to the thermal conductivity,  $\lambda_{\rm tr}^{\rm mix}$ , are given in Table 34. Notice that the results when  $X_{3p} \neq 1.00$  are not very different from the results when  $X_{3p} = 1.00$  which is, again, consistent with the assumption  $^{54}$  that the contribution to the transport properties from low lying excited electronic states is similar to the contribution from the ground state. IV. TRANSPORT PROPERTIES AT THE MIXING BOUNDARIES

During entry of a probe into the Jovian atmosphere, mixing of the ablative species with the pure atmosphere begins  $^{30}$  at 12% of the distance from the probe to

the shock front (0.294 cm from the probe) for stagnation-point peak heating. The temperature at this inner mixing boundary is 7775°K. The mole fractions of the various species at this boundary are shown in the second column of Table 35.

Mixing of the ablative species with the pure atmosphere terminates  $^{30}$  at 22% of the distance from the probe to the shock front (0.539 cm from the probe). The temperature at this outer mixing boundary is 14,756°K. The mole fractions of the various species at this boundary are shown in the third column of Table 35. The constant pressure across the shock layer is 6.29 atmospheres.

Assume that, at the inner mixing boundary, the only species that need to be considered are C, H, and O ( $X_{total}$ =0.953). The possible two body interactions are shown in Table 36. The 0-0, C-H, C-0, and H-0 interactions have not yet been considered. Also assume that, at the outer mixing boundary, the only species that need to be considered are H, e, H<sup>+</sup>, and He ( $X_{total}$ =0.999). The possible two body interactions are shown in Table 36. All of these interactions have been considered previously.

# A. The Interaction Potentials

The calculation of the transport properties at these boundaries has been considered in some  $\det^{56,57}$  and the discussion will not be repeated in this report. However, the calculation of the 0-0, C-H, C-0, and H-0 interaction potentials will be reviewed.

Transport collision integrals for the 0-0 interaction were obtained to 15,000°K by Ym and Mason.  $^{58}$  The possible molecular states of  $^{0}$ 2 that dissociate into two ground state ( $^{3}$ P) oxygen atoms are the same as the molecular states of  $^{0}$ 2 that dissociate into two ground state carbon atoms, given in Table 20. For some of the  $^{0}$ 2 bound states, RKR and/or Hulburt-Hirschfelder results  $^{51}$ ,  $^{59}$ 9 were best fit  $^{58}$ 9 with empirical potentials for which the transport collision integrals are tabulated. For other bound states and the repulsive states, the perfect pairing method was used  $^{51}$  to obtain the interaction potentials which were then best fit  $^{58}$ 9 with empirical potentials. The empirical potential parameters  $^{58}$ 8 are shown in Table 37.

The transport collision integrals obtained by Yun and Mason<sup>58</sup> are given in Table 38. Values above 15,000°K have been extrapolated from a plot of Ln (collision integral) versus Ln (T).

Now consider the C-H interaction. The states of CH that dissociate into ground state carbon and hydrogen atoms are  $^{32}$  listed in Table 39. Spectroscopic information is available  $^{32,35}$  for the  $^2\Pi$  and  $^2\Sigma$  bound states and theoretical information is available for the  $^4\Sigma$  bound state  $^{60}$  and for the repulsive  $^4\Pi$  state.  $^{61}$ 

The Morse potential was best fit to the Hulburt-Hirschfelder curve for the  $^2\Pi$  state and to the theoretical calculations  $^{60}$  for the  $^4\Sigma$  and  $^2\Sigma$  states. The exponential repulsive potential was best fit to the theoretical calculations  $^{61}$  for the  $^4\Pi$  state. The resulting parameters are shown in Table 39. Transport collision integrals were obtained for each state and then averaged according to their degeneracies.  $^{16}$  The results are given in Table 40.

Now consider the H-O interaction. The states of OH that dissociate into ground state oxygen and hydrogen atoms are  $^{32}$  listed in Table 39. Spectroscopic constants are available  $^{32,35}$  for the  $^2\pi$  and  $^2\Sigma$  bound states but neither experimental nor theoretical information is available for the  $^4\pi$  and  $^4\Sigma$  repulsive states. Thus they have been ignored in the calculation, a serious source of error. However, potential energy curves for these states can be estimated by using the perfect pairing method.

The Morse potential was best fit to the Hulburt-Hirschfelder results for the  $^2\Pi$  and  $^2\Sigma$  states. The resulting parameters are shown in Table 39. Transport collision integrals were obtained for each state and then averaged according to their degeneracies.  $^{16}$  The results are given in Table 41.

Now consider the C-O interaction. The states of CO that dissociate into ground state carbon and oxygen atoms are  $^{32}$  listed in Table 42. The spectroscopic constants are known  $^{32,35}$  for the lowest lying bound state, the  $^{1}\Sigma^{+}$  state. The next lowest lying bound state of CO, the  $^{3}\Pi_{r}$  state, does not dissociate into ground state atoms.  $^{62}$  Thus the other states listed in Table 42 have been ignored in the calculation.

The Morse potential was best fit to the Hulburt-Hirschfelder results for the  $^{1}\Sigma^{+}$  state and the transport collision integrals were calculated. The results are given in Table 43.

Using the results discussed above, the transport properties at the mixing boundaries can be calculated.  $^{56,57}$ 

# B. Errors in the Interaction Potentials

A major improvement in the results would be otained by calculating the transport properties for the Hulburt-Hirschfelder potential, using the recently developed program, for those interactions for which the necessary spectroscopic constants are available.

The major source of error in the 0-0 calculations is due to the use of the relatively crude perfect pairing method to obtain interaction potentials for more than half the states. However, these results can probably be improved since Schaefer and Harris  $^{63}$  have performed ab initio quantum mechanical calculations for 62 low lying states of  $0_2$ . Their results for the repulsive states in Table 37 can be curve fit with empirical potentials. The results should be considerably more reliable than the results obtained using the perfect pairing method.

Half the states have been ignored in the H-O calculation, clearly a major source of error. Crude estimates of the potential energy curves for the omitted states can be obtained by using the perfect pairing method. It is also desirable to have an accurabe estimate of the transport collision integrals for the H-O interaction since OH plays an important role in atmospheric photochemistry and photochemical smog. 64

All of the states of CO listed in Table 42, as well as the  $^3\Pi_r$  state, which dissociates into excited atoms, should be included in the C-O calculation. Spectroscopic information is available  $^{32}$  for the  $^1\Pi$ ,  $^1\Sigma_2^+$ , and  $^3\Pi_r$  states. Potential energy curves for the other states can be estimated using the perfect pairing method.

### V. SOME ATOM-MOLECULE AND MOLECULE-MOLECULE INTERACTIONS

Transport properties near the surface of the probe will be determined primarily by the species near the surface. The mole fractions of the species at the surface of the probe, where the temperature is 4268°K, are given<sup>30</sup> in Table 44. The species are all ablation products and most of them are diatomic or polyatomic species. Thus atom-molecule and molecule-molecule interactions must be considered as well as atomatom interactions, the only type of interaction considered previously.

## A. The Interaction Potentials

The determination of interaction potentials for atom-molecule interactions is usually difficult and the available methods are relatively crude. The problems are even greater for molecule-molecule interactions.

However, the CO-CO interaction has been considered in some detail. Mason and  ${\rm Rice}^{65} \ {\rm assumed} \ {\rm that} \ {\rm the} \ {\rm interaction} \ {\rm can} \ {\rm be} \ {\rm described} \ {\rm by} \ {\rm the} \ {\rm exponential-six} \ {\rm potential},$  using the parameters

$$\alpha = 17.0$$
  $r_{p} = 3.937 \mathring{A}$   $\epsilon/k = 119.1 ^{\circ} K$ 

The resulting transport collision integrals are given in Table 45. Since Mason and Rice<sup>65</sup> determined the parameters by comparison with experimental data on viscosities, the results in Table 45 should be quite accurate.

The CO-CO interaction is one of the few interactions involving molecules for which such a relatively straightforward procedure is available. The  $\text{He-C}_2\text{H}$  interaction will be used to illustrate a method (called the peripheral force method) for calculating interaction potential energies for interactions involving molecules. This method incorporates the assumptions that the centers of force in a molecule are located at the nucleus of each atom and that all atoms can be treated as independent entities. The name "peripheral" derives from the assumption that atoms "hidden" in the interior of molecules are not involved in the intermolecular interactions; i.e. for the  $\text{He-CH}_4$  interaction, resonable agreement with experiment is obtained by assuming that that there is no He-C interaction. The peripheral force model has been developed primarily for inverse power repulsive potentials.

The  $\text{He-C}_2\text{H}$  interaction will be used to illustrate the method since it can be assumed that the two body interactions are inverse power repulsive for this system and this system illustrates how the method can be applied to heteronuclear linear

triatomic molecules, since  $C_2H$  is linear  $^{68.69}$  with the geometry  $C^1-C^2-H^3$  where the superscripts 1, 2, and 3 label the atoms. Two "extremes" are possible during  $He-C_2H$  collisions; (1) it can be assumed that there are twice as many C-He collisions as H-He collisions which probably overcounts C-He collisions, and (2) it can be assumed that He rarely collides with the "center" carbon atom,  $C^2$  (an assumption consistent with the peripheral atom assumption); i.e. the number of C-He and H-He collisions are the same. Case 2 probably undercounts C-He collisions but, perhaps, not by much since it has been shown  $^{70}$  that, for Ar-CO $_2$  collisions (CO $_2$  has the linear geometry O-C-O), the Ar-C collisions are negligible. Only the first extreme case will be considered since the potentials used are quite crude and this case better illustrates the application of the peripheral force method to triatomic molecules.

The coordinate system used for the He- $C_2$ H interaction is given in Figure 1. The symbol r denotes the distance from the helium atom to the center of geometry of the  $C_2$ H specie. The symbols  $R_1$ ,  $R_2$ , and  $R_3$  denote the distances from the helium atom to each of the three atoms in the  $C_2$ H specie. The distances  $a_1$ ,  $a_2$ , and  $a_3$  are obtained from the estimated bond lengths;  $^{35}$ 

The assumption of independent atoms implies the assumption that the potentials are additive  $^{66}$  (which is certainly not true  $^{71}$ ); i.e.

$$V(He-C_2H) = V(He-C^1) + V(He-C^2) + V(He-H^3)$$
 (7)

However, since the  $\rm C_2H$  specie is "tumbling", it is necessary to average the atom-atom potentials (and thus the atom-molecule potential) over all angles. The averaging procedure takes the form  $^{66}$ 

$$V(atom-atom)_{av} = \frac{1}{4II} \int_{0}^{\pi} V(R) 2IIs inOd\Theta$$
 (8)

For repulsive inverse power (RIP) potentials with the form

$$V(R) = \frac{K}{R^S}$$
 equation (8) becomes <sup>66</sup>

$$V(\text{atom-atom})_{aV} = \frac{1}{2} \int_{0}^{\pi} \frac{K}{R^{S}} \sin\Theta d\Theta$$

$$= V(r)\lambda(\alpha, s)$$
(10)

where use has been made of the law of cosines; i.e.

$$R = r(1 + \alpha^2 - 2\alpha\cos\theta)^{1/2}$$

and

$$V(r) = \frac{K}{r^{S}} \tag{11}$$

A1so

$$\lambda(\alpha,s) = \frac{(1+\alpha)^{s-2} - (1-\alpha)^{s-2}}{2\alpha(s-2)(1-\alpha^2)^{s-2}}$$
 (12)

and

$$\alpha = \frac{a}{r}$$

The values of a are given in Figure 1. Using equations (7) and (10), the angle averaged atom-molecule interaction is

$$V(He-C_2H)_{av} = V(r)_{He-C_1}\lambda(\alpha_1,s_1) + V(r)_{He-C_2}\lambda(\alpha_2,s_2) + V(r)_{He-H_3}\lambda(\alpha_3,s_3)$$
 (13)

The atom-atom interactions are needed. Now 71

$$V(He-H) = \frac{2.60}{p6.06} \qquad (e.v.)$$
 (14)

The He-C interaction has not been studied from the point of view of this model. However, since it is possible to approximate the Ar-C potential by using the Ar-N potential,  $^{70}$  it is not unreasonable to approximate the He-C interaction by using the He-N potential, i.e.  $^{66}$ 

$$V(He-N) \approx V(He-C) = \frac{13.7313}{R^{6.23}}$$
 (e.v.)

However, this approximation is very crude since the justification <sup>70</sup> for the Ar-C potential is based on isoelectronic structures which is not applicable in this case.

Using the potentials given by equations (14) and (15) and the results in Figure 1, the parameters to be used in the calculations are

$$\alpha_1 = \alpha_3 = \frac{1.134}{r}$$
 $\alpha_2 = \frac{0.073}{r}$ 
 $s_1 = s_2 = 6.23$ 
 $s_3 = 6.06$ 

Results for the four terms in equation (13) can now be calculated. The results for the averaged atom-atom interactions are shown in the second, third, and fourth columns

of Table 46. Notice that  $V(\text{He-C}^2)$  is small (of the order of 10% or less) compared with  $V(\text{He-C}^1)$  which is consistent with the assumption that the He-C<sup>2</sup> interaction can be ignored.

This result suggests that it should not be assumed that He-C<sup>1</sup> and He-C<sup>2</sup> interactions are equally probable. As a crude approximation, assume that the He-C<sup>2</sup> interactions are half as frequent as the He-C<sup>1</sup> interactions (an approximation "midway" between extreme cases 1 and 2). Thus equation (13) should be modified; i.e.,  $V(\text{He-C}_2\text{H})_{av} = V(r)_{\text{He-C}} \lambda(\alpha_1, s_1) + \frac{1}{2} V(r)_{\text{He-C}} 2\lambda(\alpha_2, s_2) + V(r)_{\text{He-H}} \lambda(\alpha_3, s_3) \qquad (16)$  The resulting values of the averaged atom-molecule potential are given in the fifth column of Table 46.

The results for  $V(He-C_2H)_{av}$  have been best fit with the exponential repulsive potential in order to calculate the transport collision integrals. The best fit parameters are

$$F = 56117 \text{ e.v.}$$
  $D = 5.2002 \text{ cm}^{-1}$ 

The resulting potential energy is shown in the sixth column of Table 46. Agreement with the results for  $V(He-C_2H)_{av}$  is reasonably good.

The transport collision integrals for the  $\text{He-C}_2\text{H}$  interaction are shown in Table 47. These results have been used to calculate the binary diffusion coefficient. The results are shown in the second column of Table 48. Results obtained by Esch, et al.,  $^{72}$  using a much simpler model, are also shown in Table 48.

Now consider the C- $^{\circ}_2$  interaction. It will no longer suffice to assume that all atom-atom interactions can be approximated by the repulsive inverse power potential. Results for V(atom-atom)<sub>av</sub> should be obtained for each of the states of C<sub>2</sub> listed in Table 20, using the empirical potentials and parameters given in the table. Then, using the peripheral force model, V(C-C<sub>2</sub>)<sub>av</sub> should be determined for each state from  $^{66}$ 

$$V(C-C_2)_{av} = 2V(C-C)_{av}$$
 (17)

The results obtained for each state, using equation (17), should then be best fit with an empirical potential for which transport collision integrals have been

tabulated and the collision integrals should be averaged according to their degeneracies.  $^{16}$ 

Results have been obtained for the  $^1\Sigma_g^+$  ground state of  $^2\Sigma_g^-$ . The "true" potential for this state has been best fit with the Morse potential. When the Morse potential is substituted into equation (8), the result is

$$V(C-C_2)_{av(MP)} = \frac{\sigma \varepsilon}{\sigma} e^{\frac{C}{\sigma} r_e} e^{-\frac{C}{\sigma} r_e} e^{-\frac{C}{\sigma} r_e} e^{-2mr(1+\alpha)} - e^{-2nr(1-\alpha)} - \frac{\sigma}{4cr} e^{\frac{C}{\sigma} r_e} e^{-2mr} - e^{-2nr}$$

+ 
$$2\{e^{-mr}(1+\alpha) - e^{-nr}(1-\alpha)\} - \frac{2\sigma}{cr}\{e^{-mr} - e^{-nr}\}\}$$
 (18)

where

$$m = \frac{c}{\sigma} (1+\alpha)$$
  $n = \frac{c}{\sigma} (1-\alpha)$ 

Using equations (17) and (18) and the parameters given in Table 20 for the  $^1\Sigma_g^+$  state, the results shown in the second column of Table 49 are obtained. Notice that V(r) becomes large and negative at small values of r. This must be an artifact of the integration and cannot be physical. The results for  $V(C-C_2)_{av}$  have been best fit with the Morse potential. The best fit parameters (for  $r \ge 0.6000\text{\AA}$ ) are

$$\varepsilon \approx 6.93 \text{ e.v.}$$
  $r_p = 1.560 \text{Å}$   $c = 2.1677$ 

The resulting potential energy is shown in the third column of Table 49. The fit to the results in the second column is not very good except near  $r_e$  where, as before, the curve fit was optimized. The transport collision integrals are given in Table 50.

Now consider the  $\rm C_2$ - $\rm C_2$  interaction, a molecule-molecule interaction. According to the peripheral force model, the orientation averaged potential energy for the  $\rm C_2$ - $\rm C_2$  interaction is  $^{66}$ 

$$V(\text{molecule-molecule})_{av} = \frac{4}{(4\pi)^2} \int_0^{\pi} \int_{1-\alpha}^{1+\alpha} V(R) 2\pi \sin \theta_2 d\theta_2 2\pi \sin \theta_1 d\theta_1$$
 (19)

If the morse potential is used for V(R), the result is

$$V(C_2-C_2)_{av(MP)} =$$

$$\frac{\varepsilon\sigma^{3}}{4\alpha^{2}(cr)^{3}}e^{2\frac{c}{\sigma}}(r_{e}-r) = -4\frac{c}{\sigma}\alpha r \qquad (Mr+1) + e^{4\frac{c}{\sigma}\alpha r}(Nr+1) - 2\frac{c}{\sigma}r - 2$$

$$+\frac{2\varepsilon\sigma^3}{4\alpha^2(cr)^3}e^{\frac{c}{\sigma}(r_e-r)}[4+2\frac{c}{\sigma}r-e^{-2\frac{c}{\sigma}\alpha r}(Mr+2)-e^{-2\frac{c}{\sigma}\alpha r}(Nr+2)]$$

where

$$M = \frac{c}{\sigma} (1+2\alpha) \qquad N = \frac{c}{\sigma} (1-2\alpha)$$

Again, only the  $^1\Sigma_g^+$  ground state of  $C_2$  has been considered. Using the parameters given in Table 20, the results given in the second column of Table 51 are obtained. As was the case for the  $C-C_2$  interaction, the large negative values of V(r) at small values of r must be an artifact of the integration and will be ignored. The results in the second column of Table 51 were best fit with the Morse potential. The best fit parameters (for  $r \ge 1.058 \text{\AA}$ ) are

$$\varepsilon = 7.92 \text{ e.v.}$$
  $r_p = 1.867\text{Å}$   $c = 3.4329$ 

The resulting potential energy is shown in the third column of Table 51. The fit to the results in the second column is not very good except near  $r_e$  where, as before, the curve fit has been optimized. The transport collision integrals for the  $C_2$ - $C_2$  interaction are given in Table 52.

It is interesting to notice that the C-C, C-C<sub>2</sub>, and  $C_2$ - $C_2$  interactions become progressively "longer range"; i.e. the range of separations at which attractive interactions occur becomes progressively larger. This is shown explicitly in Table 53. In addition

$$\epsilon_{C_2-C_2}$$
 >  $\epsilon_{C-C_2}$  >  $\epsilon_{C-C}$ 

i.e. the "quasi-molecules"  $C_4$ ,  $C_3$ , and  $C_2$  have the following order of stability (according to these first order calculations);

$$C_4 > C_3 > C_2$$

The reason for these results is not entirely clear but one possible explanation is the relative polarizabilities of C and  $C_2$ . The specie  $C_2$  is almost certainly more

polarizable than C. Thus the long range attractive induced dipole-induced dipole forces would have the relative strengths

$$c_2 - c_2 > c - c_2 > c - c$$

This order is consistent with the calculated results.

The results in Tables 21, 50, and 52 can be used to estimate the transport properties in a mixture  $^{1,8}$  of C and C<sub>2</sub>. Some results for  $\lambda_{\rm tr}^{\rm mix}$  are given in Table 54.

# B. Errors in the Interaction Potentials

The interaction potentials for the  $He-C_2H$  interaction are quite crude and this interaction was considered primarily to illustrate the application of the peripheral force model to interactions involving linear triatomic molecules. There are two main sources of error in this calculation. First, as already mentioned, the He-C interaction is almost certainly seriously in error.

Second, there does not appear to be any a priori method for assigning probabilities to the  $\text{He-C}^1$  and  $\text{He-C}^2$  collisions. The assumption Collisions are half as frequent as  $\text{He-C}^1$  collisions seems intuitively reasonable on the basis of geometric and steric considerations but it is, of course, just a "good" guess.

The surprisingly good agreement between these results and those of Esch, et al.,  $^{72}$  shown in Table 48, is no reason for increased confidence in the results since the results of Esch, et al.  $^{72}$  are probably no more reliable than these results. They  $^{72}$  used the Lennard-Jones (6,12) potential; i.e.

$$V(r) = 4\varepsilon \left[ \left( \frac{\sigma}{r} \right)^{12} - \left( \frac{\sigma}{r} \right)^{6} \right]$$

and estimated  $\sigma$  from a plot of  $\sigma$  versus molecular weight for known species and estimated  $\varepsilon$  from a plot of  $\varepsilon$  versus molecular weight for known species. The good agreement in Table 48 is intriguing but almost certainly fortuitous.

For the C-C<sub>2</sub> and C<sub>2</sub>-C<sub>2</sub> interactions, only the  $^1\Sigma_g^+$  ground state of C<sub>2</sub> has been included in the calculation. The 17 other states listed in Table 20 should also be included in the calculation. These calculations are in progress.

The Morse potential has been assumed to be the "true" potential for the  $C-C_2$  and  $C_2-C_2$  interactions. However, as discussed previously, the Hulburt-Hirschfelder potential is much more accurate. If the Hulburt-Hirschfelder potential is used for V(R) in equation (8), the result is

$$\begin{split} &\frac{\epsilon}{\alpha(xr)^2} e^{xr} e^{e^{-pr}(pr+1)} - e^{-qr}(qr+1)] + \\ &\frac{\epsilon}{\alpha(xr)^2} e^{xr} e^{e^{-pr}(pr+1)} - e^{-qr}(qr+1)] + \\ &\frac{\epsilon}{\alpha} cx^2 e^{2xr} e^{e^{-pr}(pr+1)} - e^{-qr}(qr+1)] + \\ &\frac{\epsilon}{\alpha} cx^2 e^{2xr} e^{e^{-2pr}(qr+1)} - (1-\alpha)^5 e^{-2qr}(qr+1)] + (1+\alpha)^4 e^{-2pr} - (1-\alpha)^4 e^{-2qr}(qr+1) + r^2(qr+1) + r^2(qr$$

where

$$P = x(1+\alpha)$$
  $Q = x(1-\alpha)$   $X = \frac{\omega_e}{2r_e\sqrt{B_e\varepsilon}}$ 

Calculations of the  $C-C_2$  transport properties, using equation (21), are now in progress.

A general precaution about the use of the peripheral force model is necessary. Previous calculations using this model have been for interactions for which experimental data that can be used to check the potentials is available. The model seems to work reasonably well but it is not highly accurate. Alternative models may be more accurate. It cannot necessarily be assumed that this model can be used for all systems of interest and it has been assumed as "an article of faith" that the peripheral force model can be applied to the C-C<sub>2</sub> and C<sub>2</sub>-C<sub>2</sub> interactions.

# VI. OTHER COLLISION INTEGRALS

Some other transport collision integrals which may be useful are given in Tables 55 to 58. The He-C results have been obtained by using the potential given

by equation (15). Thus, as discussed previously, the results are probably not very accurate. Results have not been given for  $\sigma^2\Omega^{(1,1)*}$  for the C-C<sup>+</sup> interaction since this collision integral should be determined from charge transfer.

# VII. TRANSPORT PROPERTIES OF AIR

Using the recently developed program for calculating transport collision integrals for the Hulburt-Hirschfelder potential, as well as RKR results  $^{75}$  for some of the bound states of  $0_2$ , significant improvements on the previously calculated  $^{58}$  transport collision integrals of  $0_2$  are possible. In addition, the Hulburt-Hirschfelder program, RKR results  $^{38}$  for  $N_2$ , theoretical calculations,  $^{76}$  and the perfect pairing method can be used to improve on the previously calculated  $^{58}$  transport collision integrals of  $N_2$ . Using similar information and the peripheral force model, improved estimates of the  $N_2$ - $N_2$ ,  $0_2$ - $0_2$ , and  $N_2$ - $0_2$  interactions can also be obtained. Thus the transport properties of air can be re-evaluated. These results can be compared with experiment  $^{78}$  and thus provide a test of some of the new techniques described in this report.

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Table 1

Mole Fractions of Species in the Jovian Atmosphere as a Function of Temperature at 1 Atmosphere Pressure

Tx10 <sup>-3</sup> (OK)	H <sub>2</sub>	H	Не	_H <sup>+</sup>	He <sup>+</sup>	<u>      e                              </u>
1	0.8900		0.1100			
2	0.8854	0.0015	0.1099			
3	0.7596	0.1380	0.1024			
4	0.2054	0.7245	0.0702			
5	0.0203	0.9204	0.0594			
6	0.0032	0.9383	0.0584			
7	0.0008	0.9400	0.0582	0.0005		0.0005
8	0.0003	0.9372	0.0581	0.0022		0.0022
9	0.0001	0.9269	0.0578	0.0076		0.0076
10	0.0001	0.9019	0.0570	0.0205		0.0205
11		0.8525	0.0555	0.0460		0.0460
12		0.7699	0.0531	0.0885		0.0885
13		0.6522	0.0495	0.1492		0.1492
14		0.5091	0.0452	0.2228		0.2228
15		0.3629	0.0408	0.2981		0.2981
16		0.2378	0.0370	0.3625	0.0001	0.3626
17		0.1469	0.0342	0.4093	0.0002	0.4095
18		0.0885	0.0321	0.4392	0.0005	0.4397
19		0.0535	0.0302	0.4568	0.0013	0.4581
20		0.0307	0.0280	0.4666	0.0028	0.4694
21		0.0211	0.0249	0.4715	0.0055	0.4770
22		0.0140	0.0207	0.4733	0.0094	0.4827
23		0.0095	0.0158	0.4733	0.0140	0.4873
24		0.0067	0.0112	0.4726	0.0184	0.4910
25		0.0049	0.0076	0.4719	0.0219	0.4938

Table 2
Interaction Potentials in the Jovian Atmosphere

Inter- action	Poten- tial	Refer- ences (ab initio results)	Refer- ences (para- meters	Temp. (°K)	Pai	rameters	
H-H ( <sup>1</sup> Σ)	AIP AIP	13 13	16 16	low high	A=37.97 A=12.43	D=6 B=4.36	
H-H ( $^3\Sigma$ )	ER	13	16	all	F=60.42	D=3.013	
Не-Не	ES ER ER	14 15 15	14 21 21	1000 2000-10,000 >10,000	$\alpha$ =12.4 F=384.1 F=44.79	r <sub>e</sub> =3.135 D=4.502 D=2.903	ε/k=9.16
H <sub>2</sub> -H <sub>2</sub>	ES ER	14 16	14 16	≤7000 >7000	α=14.0 F=116.5	r <sub>e</sub> =3.337 D=2.859	ε/k=37.3
Н <sub>2</sub> -Н	ER	16	16	all	F=61.5	D=2.952	
H <sub>2</sub> -He	ES ER	14	<b>14</b> 8	≤3000 >3000	α=13.22 F=211.5	r <sub>e</sub> =3.244 D=3.699	ε/k=18.27
H-He	ER	15	21	all	F=74.73	D=3.159	
H-Н <sup>+</sup> (2р <sub>0</sub> )	ER	17	21	all	F=56.38	D=1.719	
$H-H^+(1s_{\sigma})$	MP	17	8	all	C=1.230	r <sub>e</sub> =1.100	ε=2.800
He-H <sup>+</sup>	MP	18	22	all	C=1.230	r <sub>e</sub> =0.762	ε=1.905
$He-He^+(^2\Sigma_u^-)$	MP	19	19	all	C=1.637	r <sub>e</sub> =1.080	ε=2.16
$He-He^+(^2\Sigma_g)$	ER	19	19	all	F=44.40	D=2.157	
$H-He^+(^1\Sigma)$	ER	20	8	all	F=149.2	D=3.019	
$H-He^+(^3\Sigma)$	ER	20	8	all	F=157.75	D=3.716	

The parameters have been chosen so that V(r) is in electron volts and r is in Angstroms.

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$
1	5.235	5.954
2	4.133	4.743
<b>3</b> ·	3.570	4.118
4	3.230	3.742
5	3.028	3.500
6	2.884	3.281
7	2.760	3.063
8	2.622	2.883
9	2.479	2.730
10	2.356	2.598
11	2.249	2.483
12	2.754	2.380
13	2.070	2.289
14	1.993	2.205
15	1.924	2.130
15.5	1.890	2.093
16	1.858	2.058
16.5	1.829	2.026
17	1.801	1.995
17.5	1.773	1.964
18	1.746	1.934
19	1.699	1.882
20	1.653	1.833
21	1.610	1.785
22	1.569	1.738
23	1.532	1.697
24	1.496	1.658
25	1.463	1.620

Table 4
Collision Integrals for the He-He Interaction

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ ( $\mathring{A}^2$ )	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$
1	3.023	3.779
2	2.508	3.034
3	2.231	2.719
4	2.050	2.507
5	1.911	2.351
6	1.803	2.227
7	1.717	2.126
8	1.637	2.035
9	1.577	1.964
10	1.517	1.894
11	1.345	1.785
12	1.277	1.700
13	1.218	1.625
14	1.160	1.552
15	1.118	1.529
15.5	1.111	1.515
16	1.104	1.501
16.5	1.086	1.481
17	1.069	1.461
17.5	1.052	1.442
18	1.036	1.423
19	1.006	1.387
20	0.978	1.354
21	0.951	1.323
22	0.926	1.294
23	0.903	1.265
24	0.880	1.239
25	0.859	1.213

$Tx10^{-3}$ (°K)	$\sigma^{2}\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$
1	5.210	6.002
2	4.367	5.328
3	3.794	4.732
4	3.419	4.293
5	3.142	3.966
6	2.925	3.710
7	2.747	3.497

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}$ (Å <sup>2</sup> )	A*	<u>B</u> *
1	4.157	5.134	1.235	1.200
2	3.270	4.100	1.254	1.222
3	2.802	3.549	1.266	1.237
4	2.492	3.180	1.276	1.249
5	2.265	2.907	1.284	1.259
6	2.088	2.694	1.290	1.268
7	1.943	2.518	1.296	1.276

Table 7 Collision Integrals for the  $\mathrm{H}_2\text{-He}$  Interaction

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$	A*	B*
1	4.059	4.829	1.189	0.911
2	3.517	4.292	1.220	0.931
3	3.244	4.041	1.246	0.937
4	2.511	3.111	1.239	1.204
5	2.329	2.899	1.245	1.211
6	2.178	2.723	1.250	1.250
7	2,062	2.587	1.254	1.223

 $\label{thm:collision} \mbox{Table 8}$  Collision Integrals for the H-He Interaction

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}$ (Å <sup>2</sup> )	_A*	<b>B</b> *
1	3.858	4.747	1.231	1.194
2	3.050	3.809	1.249	1.215
3	2.632	3.318	1.261	1.230
4	2.349	2.983	1.270	1.241
5	2.145	2.740	1.277	1.251
6	1.977	2.538	1.284	1.259
7	1.849	2.384	1.289	1.266
8	1.742	2.254	1.294	1.273
9.	1.646	2.137	1.298	1.279
10	1.560	2.033	1.303	1.285
11	1.492	1.949	1.306	1.291
12	1.425	1.867	1.310	1.296
13	1.368	1./96	1.313	1.301
14	1.318	1.734	1.316	1.305
15	1.269	1.674	1.319	1.309
15.5	1.245	1.644	1.321	1.312
16	1.222	1.615	1.322	1.314
16.5	1.202	1.589	1.323	1.316
17	1.182	1.565	1.324	1.318
17.5	1.165	1.544	1.325	1.320
18	1.149	1.524	1.326	1.321
19	1.110	1.475	1.329	1.326
20	1.078	1.435	1.331	1.329
21	1.047	1.396	1.333	1.333
22	1.016	1.357	1.336	1.336
23	0.992	1.326	1.337	1.337
24	0.968	1.296	1.339	1.342
25	0.944	1.266	1.341	1.345

Table 9 Collision Integrals for the  $\mathrm{H}\text{-H}^{^+}$  Interaction

	•			
Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}(\mathring{A}^2)$	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$	_ <u>A</u> *	
8	30.9	6.49	0.210	1.38
9	30.4	6.02	0.198	1.38
10	30.0	5.61	0.187	1.38
11	29.6	5.28	0.178	1.37
12	29.2	4.98	0.170	1.37
13	28.9	4.74	0.164	1.37
14	28.6	4.51	0.158	1.37
15	28.3	4.31	0.152	1.37
15.5	28.2	4.21	0.149	1.37
16	28.1	4.12	0.147	1.37
16.5	27.9	4.03	0.144	1.37
17	27.8	3.95	0.142	1.37
17.5	27.7	3.87	0.139	1.37
18	27.6	3.79	0.137	1.37
19	27.4	3.66	0.134	1.37
20	27.2	3.53	0.130	1.37
21	27.0	3.42	0.127	1.37
22	26.8	3.31	0.123	1.37
23	26.6	3.21	0.120	1.37
24	26.5	3.11	0.117	1.36
25	26.3	3.02	0.115	1.36

Table 10

Collision Integrals for the H-e Interaction

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,i)*}(\mathring{A}^2)$	$\sigma^2\Omega(2,2)$ * ( $\mathring{A}^2$ )	<u> </u>	<u>B</u> *
8	6.44	8.34	1.30	-1.73
9	6.13	7.92	1.29	-1.79
10	5.87	7.56	1.29	-1.84
11	5.63	7.24	1.29	-1.88
12	5.42	6.96	1.28	-1.92
13	5.24	6.71	1.28	-1.96
14	5.07	6.49	1.28	-1.99
15	4.92	5.29	1,28	-2.03
15.5	4.85	6.19	1.28	-2.04
16	4.78	6.10	1.28	-2.05
16.5	4.71	6.01	1.28	-2.07
17	4.65	5.93	1.28	-2.08
17.5	4.59	5.85	1.27	-2,10
18	4.53	5.77	1.27	-2.11
19	4.42	5.63	1.27	-2.14
20	4.32	5.50	1.27	-2.15
21	4.22	5.37	1.27	-2.18
22	4.14	5.25	1.27	-2.20
23	4.05	5.14	1.27	-2.22
24	3.97	5.04	1.27	-2.25
25	3.90	4.94	1.27	-2.26

Table ]] Collision Integrals for the  $\operatorname{He-H}^{+}$  Interaction

	corrision integra	is for the ne-ti	incer action	
$Tx10^{-3}$ (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}$ (Å <sup>2</sup> )	A*	
8	1.923	2.214	1.207	1.432
9	1.706	2.014	1.181	1.422
10	1.535	1.842	1.200	1.412
11	1.397	1.699	1.216	1.403
12	1.280	1.577	1.233	1.394
13	1.181	1.474	1.248	1.386
14	1.098	1.384	1.261	1.379
15	1.024	1.305	1.275	1.372
15.5	0.9906	1.269	1.282	1.369
16	0.9593	1.235	1.288	1.366
16.5	0.9307	1.203	1.293	1.363
17	0.9037	1.173	1.298	1.360
17.5	0.8777	1.114	1.304	1.357
18	0.8537	1.117	1.310	1.355
19	0.8079	1.067	1.321	1.350
20	0.7680	1.022	1.331	1.345
21	0.7312	0.9802	1.340	1.340
22	0.6977	0.9422	1.350	1.337
23	0.6678	0.9072	1.359	1.341
24	0.6397	0.8752	1.368	1.345
25	0.6141	0.8450	1.376	1.349

Table 12

Collision Integrals for the He-e Interaction

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )
8	2.625
9	2.629
10	2.632
11	2.632
12	2.630
13	2.627
14	2.623
15	2.620
15.5 16 16.5 17	2.617 2.615 2.613 2.611
17.5	2.608
18	2.606
19	2.595
20	2.585
21	2.578
22	2.567
23	2.554
24	2.544
25	2.483
•	$= \sigma^2 \Omega(1,1)$ *
A* = B* =	l for all values of T

,

Table 13
Collision Integrals for the  $H^{\dagger}-H^{\dagger}$ ,  $He^{\dagger}-He^{\dagger}$ , and  $H^{\dagger}-He^{\dagger}$  Interaction

Collision Int	egrals for the $ extsf{H}^{ extsf{T}}$ -		H <sup>-</sup> -He <sup>-</sup> Inte	ractions
Tx10 <sup>-3</sup> (°K)	$\lambda_{d}^{2\Omega}^{(1,1)*}$ (Å <sup>2</sup> )	$\lambda_{d}^{2\Omega}(2,2)* (\mathring{A}^{2})$	A*	B*
8	845.11	960.46	1.1365	1.1735
9	596.71	685.89	1.1495	1.1892
10	445.79	517.35	1.1605	1.2026
11	343.68	402.01	1.1698	1.2136
12	272.11	320.73	1.1765	1.2215
13	224.96	266.37	1.1809	1.2268
14	191.57	226.67	1.1832	1.2295
15	165.37	195.75	1.1837	1.2301
15.5	155.96	184.50	1.1830	1.2292
16	147.37	174.23	1.1823	1.2284
16.5	139.88	165.19	1.1809	1.2268
17	132.98	156.86	1.1796	1.2252
17.5	127.17	149.78	1.1778	1.2231
18	121.77	143.21	1.1760	1.2210
19	113.26	132.74	1.1720	1.2162
20	105.03	122.67	1.1680	1.2115
21	98.400	114.54	1.1641	1.2069
22	91.607	106.31	1.1605	1.2026
23	85.709	99.192	1.1574	1.1988
24	81.103	93.610	1.1542	1.1950
25	76.319	87.863	1.1513	1.1914

The values of A\* and B\* can be taken to be unity at all values of T for the  $H^+-H^+$  and  $He^+-He^+$  interactions.

Table 14
Collision Integrals for the e-e Interaction

Tx10 <sup>-3</sup> (°K)	$\lambda_{d}^{2\Omega}(1,1)*(\mathring{A}^{2})$	$\lambda_{d}^{2}\Omega^{(2,2)*}$ (Å <sup>2</sup> )
8	841.70	952.55
9	593.73	678.84
10	443.11	510.91
11	341.26	396.01
12	269.90	315.28
13	222.91	261.27
14	189.67	221.88
15	163.61	191.22
15.5	154.27	180.11
16	145.73	169.96
16.5	138.29	161.04
17	131.44	152.83
17.5	125.68	145.86
18	120.32	139.40
19	111.89	129.11
20	103.74	119.22
21	97.174	111.25
22	90.438	103.17
23	84.595	96.190
24	80.036	90.718
25	75.302	85.097

Table 15 Collision Integrals for the  $H^{\dagger}-e$  and  $He^{\dagger}-e$  Interactions

Tx10 <sup>-3</sup> (°K)	$\frac{\lambda_{d}^{2} \Omega^{(1,1)*} (\mathring{A}^{2})}{}$	$\frac{\lambda_d^2 \Omega^{(2,2)*} (\mathring{A}^2)}{}$	A* 	B*
8	881.11	967.49	1.0980	1.2866
9	631.96	696.15	1.1016	1.3030
10	478.96	528.84	1.1041	1.3168
11	374.08	413.57	1.1056	1.3300
12	299.66	331.47	1.1061	1.3384
13	249.52	276.10	1.1065	1.3422
14	212.58	235.21	1.1065	1.3433
15	183.61	203.12	1.1062	1.3458
15.5	172.50	191.29	1.1089	1.3484
16	162.39	180.49	1.1115	1.3510
16.5	153.78	170.89	1.1113	1.3499
17	145.87	162.07	1.1111	1.3488
17.5	139.09	154.50	1.1108	1.3467
18	132.80	147.48	1.1106	1.3446
19	122.72	136.22	1.1100	1.3385
20	113.07	125.44	1.1094	1.3337
21	105.30	116.75	1.1087	1.3278
22	97.519	108.04	1.1079	1.3245
23	90.819	100.55	1.1071	1.3224
24	85.546	94.648	1.1064	1.3179
25	80.173	88.640	1.1056	1.3142

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}$ (Å <sup>2</sup> )	A*	
15	1.483	1.904	1.283	1.266
15.5	1.461	1.877	1.284	1.268
16	1.440	1.852	1.286	1.270
16.5	1.420	1.828	1.287	1.272
17	1.401	1.805	1.288	1.274
17.5	1.382	1.782	1.289	1.275
18	1.364	1.760	1.290	1.277
19	1.329	1.719	1.293	1.280
20	1.298	1.680	1.294	1.283
21	1.266	1.642	1.297	1.286
22	1.237	1.606	1.298	1.288
23	1.210	1.573	1.300	1.291
24	1.183	1.539	1.301	1.294
25	1.158	1.510	1.304	1.296

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}(\mathring{A}^2)$	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$	A*_	<u> </u>
15	91.7	2.69	0.030	1.37
15.5	91.9	2.63	0.029	1.37
16	92.1	2.57	0.029	1.37
16.5	92.2	2.52	0.028	1.37
17	92.4	2.47	0.027	1.37
17.5	92.6	2.42	0.027	1.37
18	92.9	2.37	0.026	1.37
19	93.2	2.29	0.025	1.38
20	93.4	2.22	0.024	1.38
21	93.6	2.14	0.023	1.38
22	93.9	2.07	0.022	1.38
23	94.1	2.01	0.021	1.38
24	94.3	1.96	0.021	1.38
25	94.5	1.90	0.020	1.38

Table 18 The  $\mbox{He-H}^{\dagger}$  Interaction Potential as a Function of Internuclear Separation

r (Å)	V <sub>ab inito</sub> (ev) <sup>22</sup>	V <sub>Morse</sub> (ev)
0.5290	0.29	-0.69
0.6348	-1.54	-1.63
0 7406	-1.90	-1.90
0.8464	-1.84	-1.83
0.9522	-1.66	-1.63
1.1638	-1.28	-1.13
1.2696	-1.12	-0.91
1.3754	-0.99	-0.72
1.4812	-0.88	-0.57
1.5870	-0.78	-0.45
1.6928	-0.48	-0.35
1.7986	-0.36	-0.27
1.9044	-0.27	-0.21
2.0102	-0.20	-0.16
2.1160	-0.16	-0.12
2.2218	-0.12	-0.09
2.3276	-0.10	-0.07

Table 19

Mole Fractions of Neutral Carbon Species as a Function of Temperature

and Distance From the Stagnation Point

		*		
Distance (cm)	T (°K)	x (c)	x (c <sub>2</sub> )	$\frac{x(c_3)}{x(c_3)}$
0.00	4268	0.0258	0.0275	0.3034
0.05	4633	0.0724	0.0558	0.3368
0.10	5141	0.2075	0.0971	0.2631
0.15	5553	0.3870	0.1175	0.1466
0.20	5855	0.5236	0.1211	0.0487
0.24	6236	0.6367	0.0937	0.0081
0.27	6601	0.6662	0.0571	0.0013
0.29	7775	0.6534	0.0098	
0.32	9790	0.5006	0.0008	
0.33	10,742	0.3992	0.0002	
0.34	11,424	0.3076	0.0001	
0.35	11,980	0.2300		
0.37	12,456	0.1692		
0.38	12,821	0.1217		
0.39	13,170	0.0857		
0.40	13,392	0.0609		
0.42	13,598	0.0429		
0.43	13,773	0.0302		
0.44	13,931	0.0211		
0.45	14,058	0.0147		
0.46	14,169	0.0102		
0.49	14,344	0.0046		
0.51	14,487	0.0018		
0.54	14,614	0.0006		

State	<u>Potential</u>	Par	rameters*
$^{1}\Sigma_{g}^{+}$	MP	C=2.500	r <sub>e</sub> =1.2420 ε=6.36
ำ ก <sub>น</sub>	MP	C=2.193	r <sub>e</sub> =1.3119 ε=6.28
$3_{\Sigma_{\mathbf{g}}^{-}}$	MP	C 2.182	$r_e^{=1.3692}$ $\epsilon = 5.59$
$^{1}\pi_{u}$	MP	C=2.339	$r_e = 1.3184$ $\epsilon = 5.33$
$3_{\Sigma_{\mathbf{u}}^{+}}$	MP	C=2.997	$r_e = 1.23$ $\epsilon = 4.70$
1 <sub>Ag</sub>	MP	C=2.697	$r_e^{=1.39}$ $\varepsilon = 4.45$
$^{1}\Sigma_{g}^{+}$ 2	MP	C=2.238	$r_e = 1.38$ $\epsilon = 4.03$
$^{3}$ n $_{g}$	MP	C=2.978	$r_e = 1.2660$ $\epsilon = 3.89$
5 <sub>II</sub> g	MP	C=2.739	$r_e = 1.46$ $\epsilon = 3.54$
$^{5}\Sigma_{g}^{+}$	MP	C=4.044	$r_e = 1.35$ $\epsilon = 2.55$
1 <sub>π</sub> g	MP	C=4.328	r <sub>e</sub> =1.2552 ε =2.11
$^{3}$ u	MP	C=4.071	$r_e$ =1.51 $\epsilon$ =2.02
$^{1}\Sigma_{\mathbf{u}}^{-}$	MP	C=3.474	$r_e$ =1.90 $\epsilon$ =1.98
5 <sub>Σ</sub> -	ER	F=309.5	D=0.4060
5 <sub>II</sub> u	ER	F=400.0	D=0.3601
$^{5}_{\Delta_{f g}}$	ER	F=504.3	D=0.3931
<sup>5</sup> π <sub>u</sub> <sup>5</sup> Δ <sub>g</sub> <sup>3</sup> Σ <sup>+</sup> <sub>u</sub> 2 <sup>5</sup> Σ <sup>+</sup> <sub>α</sub> 2			
$^{5}\Sigma_{g}^{+}$ 2			

<sup>\*</sup>The parameters have been chosen so that V(r) is in electron volts and r is in Angstroms.

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Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$
1	10.4254	10.9550
2	8.9651	9.3214
3	8.1203	8.4021
4	7.5021	7.7863
5	7.0400	7.3393
6	6.6325	6.9795
7	6.2860	6.6854
8	5.9865	6.4239
9	5.6957	6.1864
10	5.4336	5.9610
11	5.2017	5.7595
12	4.9823	5.5740
13	4.7842	5.3833
14	4.6072	5.2186
15	4.4449	5.0620
16	4.2958	4.9137
17	4.1499	4.7715
18	4.0144	4.6372
19	3.8860	4.4885
20	3.7654	4.3663
21	3.6374	4.2454
22	3.5324	4.1436
23	3.4366	4.0423
24	3.3474	3.9462
25	3.2630	3.8553

Table 22 Potential Energy Curves for the Hulburt-Hirschfelder Potential and the Morse Potential for the  $^1\Sigma_g^+$  State of C2

r (A)	V <sub>HH</sub> (ev)	V <sub>Morse</sub> (eV)
1.058	-4.42	-4.83
1.111	-5.49	-5.67
1.242	-6.36	-6.36
1.323	-6.15	-6.19
1.429	-5.47	-5.65
1.588	-4.15	-4.59
1.852	-2.26	-2.94
2.117	-1.13	-1.76
2.646	-0.29	-0.59
3.704	-0.03	-0.06
5.292	0	0

Table 23 The Hulburt-Hirschfelder Potential for the  $^1\pi_g$  State of  $\mathrm{C}_2$ 

<u>r/o</u>	$V(r)/\varepsilon$
1.150	-0.996
1.300	-0.560
1.450	-0.115
1.600	0.036
1.750	0.042
1.900	0.021
2.050	0.007
2.200	0.001
2.350	0
2.500	-0.001
2.650	0

Table 24 Comparison of the Collision Integrals for the  $^1 \pi_g^+$  State of C $_2$  for the Hulburt-Hirschfelder and Morse Potentials

	TOT CHE HATDUTC	-IIII Scille Laci ali		13
$T \times 10^{-3}$ (°K)	$\sigma^2\Omega^{(1,1)*}$ (HH)	$\sigma^2\Omega(1,1)*$ (MP)	$\sigma^{2}\Omega^{(2,2)*}$ (HH)	$\sigma^{2}\Omega^{(2,2)*}$ (MP)
1	7.4841	11.6132	7.8147	11.5269
2	6.4250	10.1956	6.5837	9.7526
3	5.8522	9.3141	5.9310	8.8155
4	4.6151	8.7405	4.6097	8.2219
5	4.4527	8.2450	4.4743	7.7398
6	4.3244	7.8490	4.3666	7.3881
7	4.2186	7.5016	4.2777	7.0946
8	4.1288	7.2046	4.2017	6.8832
9	4.0513	6.8877	4.1359	6.6577
10	3.9835	6.6343	4.0782	6.4829
11	3.9231	6.3719	4.0267	6.3035
12	3.8686	6.1279	3.9801	6.1357
13	3.8192	5.9332	3.9378	5.9996
14	3.7735	5.7057	3.8985	5.8370
15	3.7144	5.5028	3.8505	5. <sup>-</sup> 879
16	3.6084	5.3264	3.7702	5.5547
17	3.5102	5.1482	3.6951	5.4163
18	3.4208	4.9687	3.6264	5.2730
19	3.3382	4.8185	3.5624	5.1498
20	3.2620	4.6680	3.5030	5.0234
21	3.1905	4.5177	3.4469	4.8940
22	3.1233	4.3678	3.3939	4.7620
23	3.0480	4.2188	3.3284	4.6277
24	2.9740	4.1004	3.2630	4.5188
25	2.9069	3.9829	3.2033	4.4090

The collision integrals are given ir units of  $(Angstroms)^2$ .

Table 25

RKR and Hulburt-Hirschfelder Potential Energy Results

for the $^1\Sigma_{f q}^+$ , $^3\pi_{f u}$ , and $^1\pi_{f q}$ States of $^0\Sigma_2$								
	$^{1}\Sigma_{g}^{+}$		9 <b>-</b>	$3_{\Pi_{\mathbf{u}}}$	_		1 <sub>ng</sub>	
<u>r (Å)</u>	V (RKR)	V (HH)	r(Å)	V (RKR)	V (HH)	r(Å)	V (RKR)	V (HH)
1.134	0.564	0.564	1.098	2.149	2.164	1.103	1.201	1.377
1.156	0.341	0.342	1.105	1.977	1.993	1.111	1.078	1.187
1.190	0.115	0.117	1.113	1.802	1.808	1.120	0.922	1.001
1.301	0.115	0.115	1.121	1 525	1.635	1.132	0.740	0.790
1.349	0.341	0.344	1.131	1.445	1.435	1.148	0.541	0.563
1.384	0.564	0.564	1.141	1.262	1.253	1.169	0.330	0.341
			1.152	1.076	1.070	1.202	0.112	0.119
			1.164	0.886	0.891	1.313	0.112	0.109
			1.179	0.695	0.696	1.365	0.330	0.344
			1.197	0.500	0.500	1.405	0.540	0.575
			1.221	0.302	0.297	1.444	0.740	0.815
			1.257	0.101	0.100	1.488	0.922	1.083
			1.374	0.101	0.101	1.545	1.078	1.398
			1.425	0.302	0.303	1.621	1.201	1.729
			1.463	0.500	0.501			
			1.495	0.695	0.692			
			1.525	0.886	0.852			
			1.553	1.076	1.073			
			1.580	1.262	1.260			
			1.605	1.445	1.437			
			1.630	1.625	1.615			
			1.655	1.802	1.793			
			1.679	1.977	1.964			
			1.702	2.149	2.127			

Potential energy is given in electron volts.

Table 26 Molecular Orbital Arrangements for the  $^3\Sigma_{u2}^+$ ,  $^5\pi_{g2}^+$ ,  $^3\Sigma_u$ , and  $^3\Sigma_g^-$  States of  $^2$  According to the Perfect Pairing Method

Molecular Orbital	3 <sub>Σ</sub> + 2	5 <sub>Σ</sub> + g2	$\frac{3}{\Pi_{\mathbf{u}}}$	$\frac{3}{\Sigma}g$
თ <b>*2p</b>	†			
л <b>*-</b> 2р		<b>†</b>		
π <sub>g</sub> 2p π <sub>g</sub> *+2p		<b>†</b>		
π_2p	+	<b>†</b>	†	<b>†</b>
п <mark>и</mark> 2р	+	<b>†</b>	<b>†</b> ‡	<b>†</b>
σ <sub>g</sub> 2p	<b>†</b>		<b>↑</b>	<b>†</b> +

For each state, eight electrons fill the molecular orbitals  $(\sigma_g \text{ls})^2 (\sigma_u^* \text{ls})^2 (\sigma_g 2 \text{s})^2 (\sigma_u^* 2 \text{s})^2$ 

Potential Energy Curves for the  $^3\Sigma^+_{u2}$  and  $^5\Sigma^+_{g2}$  States of C $_2$  Obtained by the Perfect Pairing Method

r (Å)	$V (^{3}\Sigma_{u2}^{+})$	$V (^{5}\Sigma_{g2}^{+})$
1.80		3.38
1.85		3.04
1.90		2.70
1.95		2.40
2.00		2.14
2.05		1.90
2.10	0.44	1.68
2.15	0.42	1.50
2.20	0.42	1.32
2.25	0.42	1.16
2.30	0.40	1.02
2.35	0.38	0.90
2.40	0.36	0.80
2.45	0.36	0.70
2.50	0.34	0.62
2.55	0.30	0.56
2.60	0.28	0.50
2.65	0.26	0.44
2.70	0.24	0.38
2.75	0.22	0.34
2.80	0.22	0.30
2.85	0.20	0.26
2.90	0.18	0.24
2.95	0.16	0.22
3.00	0.16	0.18

The potential energy is in electron volts.

Table 28

Mole Fraction of Carbon Atoms in the Three Lowest Electronic States

Tx10 <sup>-3</sup> (°K)	<u>X (<sup>3</sup>P)</u>	<u>x (<sup>1</sup>D)</u>	x (1s)
1	1.00	0.00	0.00
2	1.00	0.00	0.00
3	1.00	0.00	0.00
4	0.99	0.01	0.00
5	0.97	0.03	0.00
6	0.95	0.05	0.00
7	0.94	0.06	0.00
8	0.92	0.08	0.00
9	0.90	0.10	0.00
10	0.88	0.11	0.00
11	0.86	0.13	0.01
12	0.84	0.14	0.01
13	0.82	0.15	0.01
14	0.80	0.16	0.01
15	0.78	0.16	0.01
16	0.76	0.17	0.01
17	0.73	0.17	0.01
18	0.71	0.17	0.01
19	0.68	0.17	0.02
20	0.65	0.17	0.02
21	0.61	0.17	0.02
22	0.58	0.17	0.02
23	0.55	0.16	0.02
24	0.52	0.16	0.02
25	0.49	0.15	0.02

Table 29 Spectroscopic Constants for the States of  $\rm C_2$  that Dissociate into a  $\rm ^3P$  Carbon Atom and a  $\rm ^1D$  Carbon Atom

		a P C	arbon Atom	and a	'D Carbon At	Om	
State	$\frac{g_{i}}{g_{i}}$	ε (e.v.)	r <sub>e</sub> (Å)	$\frac{\omega_{\mathbf{e}}}{}$	$\frac{\omega_{\mathbf{e}}^{\chi}\mathbf{e}}{}$	<u>B</u> e	$\frac{\alpha_{e}}{e}$
3 <sub>∲</sub> g	6	2.88	1.53	1290	9.0	1.20	0.011
3 <sub>II</sub> a2	6	2.56	1.535	1107	39.26	1.1922	0.0242
$\Delta_{0.12}$	6	2.55	1.51	1380	9.2	1.23	0.011
JI -	5	2.35	1.49	1340	9.5	1.265	0.012
Σ	3	2.25	1.44	1660	10.2	1.355	0.011
JI	6	2.08	2.31	1000			
<b>Φ</b> ,,	6						
3 <sub>0</sub>	6						
3 <sub>Δ</sub> <sub>g</sub> 3 <sub>Σ</sub> <sup>+</sup> <sub>g</sub>	3						
3 <sub>Σ</sub> -	3						
30,13	6						
$\Delta_{\Delta_{02}}$	6						
JI3	6						
IIaa	6						
II	6						
$\Sigma_{\Sigma}^{\alpha}$	3						
$\Sigma_{12}$	3						
3 <sub>Σ</sub> -	3						

The constants  $\omega_e$ ,  $\omega_e \chi_e$ ,  $B_e$  and  $\alpha_e$  are in cm<sup>-1</sup>.

Table 30 Interaction Potential Parameters for the  $C(^3P)-C(^1D)$  Interaction

State	<u>Potential</u>	Parameters		
3 <sub>p</sub> g	MP	C=3.1319		
3 <sub>∏</sub> g2	MP	C=2.6543		
3 <sub>∆</sub> u2	MP	C=4.4409		
$^{3}_{ m II}$ g3	MP	C=3.4789		
3 <sub>2</sub> +	MР	C=6.2189		
3 <sub>II</sub> u2	MP	C=2.3878		
$^{3}\Phi_{u}$	MP	C=4.0000	r <sub>e</sub> =3.5827	ε=0.03
$^{3}\Delta_{g}$	MP	C=4.6451	r <sub>e</sub> =3.5827	ε=0.01
3 <sub>Δ</sub> g 3 <sub>Σ</sub> + g	ER	F=66.35	D=1.0235	
$3_{\Sigma_{\mathbf{u}}^{-}}$	MP	C=2.8790	r <sub>e</sub> =2.6460	ε=0.14
$^3\Delta$ u $^3$				
3 <sub>∆</sub> g2				
<sup>3</sup> п u3				
3 <sub>II</sub> g4				
3 <sub>II</sub> <sub>u4</sub>				
$3_{\Sigma_{\mathbf{g}}}$				
3 <sub>Σ</sub> - g 3 <sub>Σ</sub> - u <sub>2</sub> 3 <sub>Σ</sub> - g <sub>2</sub>				
3 <sub>Σ</sub> -				

The parameters have been chosen so that V(r) is in electron volts and  ${\bf r}$  is in Angstroms.

Table 31  $\label{eq:Table 31}$  Transport Collision Integrals for the C(  $^3\text{P}$  )-C(  $^1\text{D}$  ) Interaction

$T \times 10^{-3} ('K)$	$\sigma^2\Omega^{(1,1)*}(\mathring{A}^2)$	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$	A*	<b>B</b> *
1	13.7140	14.3641	1.0556	1.1473
2	11.3329	11.8458	1.0670	1.1661
3	10.0213	10.5703	1.0719	1.1919
4	9.0667	9.7169	1.0830	1.2247
5	8.2998	9.0545	1.0969	1.2525
6	7.6588	8.4959	1.1115	1.2768
7	7.1248	8.0121	1.1252	1.2953
8	6.6458	7.5643	1.1366	1.3058
9	6.2377	7.1695	1.1465	1.3121
10	5.8680	6.7992	1.1550	1.3147
11	5.5559	6.4815	1.1618	1.3150
12	5.2661	6.1777	1.1679	1.3131
13	5.0105	5.9048	1.1727	1.3104
14	4.7796	5.6547	1.1768	1.3065
15	4.5777	5.4332	1.1802	1.3026
16	4.3820	5.2157	1.1834	1.2987
17	4.2133	4.9193	1.1859	1.2946
18	4.0600	4.8537	1.1880	1.2905
19	3.9160	4.6999	1.1898	1.2873
20	3.7889	4.5450	1.1913	1.2825
21	3.6662	4.4043	1.1927	1.2790
22	3.5570	4.2789	1.1939	1.2762
23	3.4560	4.1620	1.1952	1.2729
24	3.3635	4.0545	1.1961	1.2699
25	3.2712	3.9530	1.1970	1.2672

Table 32 Spectroscopic Constants and Interaction Potential Parameters for the  $C(^1D)\text{-}C(^1D) \ \ Interaction$ 

State	<u>'s</u>	ε ( <b>e.v.</b> )	r <sub>e (Å)</sub>	Morse Para	ameters	
l <sub>p</sub> g	2	3.56	1.51	C=2.7084		
1 <sub>π</sub> g2	2	2.60	1.46	C=3.3321		
1 <sub>Σ</sub> - u2	1	2.49	1.45	C=4.5192		
$^{1}\Delta_{u}$	2	2.35	1.39	C=3.7762		
1 <sub>II</sub> g3	2	2.07	1.45	C=4.8855		
۱ <sub>۲g</sub>	2			C=4.3224	r <sub>e</sub> =3.5825 e	=0.02
1 <sub>Ф</sub> u	2			C=4.3224	r <sub>e</sub> =3.5825 e	=0.02
$^{1}\Delta_{g}$	2					
$^{1}\Delta_{g2}$	2					
$^{1}\Sigma_{g}^{+}$	1					
<sup>1</sup> Σ <sup>+</sup> g2	1					
$^{1}\Sigma_{g3}^{+}$	1					
1 <sub>nu</sub>	2					
1 <sub>πu2</sub>	2					
1 <sub>Σ</sub> -	1					

The parameters have been chosen so that V(r) is in electron volts and r is in Angstroms.

Table 33  $\label{eq:Table 33}$  Transport Collision Integrals for the C( $^1\text{D})\text{-C(}^1\text{D})$  Interaction

transport contracts	-	
Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ ( $^2$ )	$\sigma^2\Omega^{(2,2)*}$ (Å <sup>2</sup> )
	8.3971	8.7779
1	6.8097	7.2654
2	5.8666	6.3915
3	5.2095	5.7754
4	4.7130	5.3022
5	4.3273	4.9241
6		4.6095
7	4.0144	4.3456
8	3.7558	4.1189
9	3.5383	3.9293
10	3.3595	3.7598
11	3.2011	3.6151
12	3.0675	3.4860
13	2.9490	3.3701
14	2.8433	3.2661
15	2.7502	
16	2.6686	3.1772
17	2.5945	3.0952
18	2.5261	3.0190
19	2.4633	2.9492
20	2.4080	2.8881
21	2.3552	2.8293
22	2.3088	2.7773
23	2.2633	2.7266
	2.2203	2.6784
24	2.1826	2.6363
25	_ : , _ : :	

Table 34 The Translational Contribution to the Thermal Conductivity,  $\lambda_{tr}^{mix}$  (10 W/m/°K), for a Mixture of  $^3P$  and  $^1D$  Carbon Atoms

X=0.00	X=0.25	X=0.50	X=0.75	X=1.00
0.87	0.70	0.64	0.63	0.69
1.48	1.19	1.08	1.06	1.15
2.06	1.64	1.49	1.45	1.57
2.63	2.08	1.87	1.81	1.95
3.21	2.52	2.25	2.16	2.31
3.78	2.95	2.62	2.51	2.67
4.36	3.39	2.99	2.85	3.01
4.95	3.84	3.37	3.20	3.34
5.54	4.29	3.75	3.55	3.68
6.12	4.75	4.15	3.91	4.03
6.71	5.21	4.54	4.27	4.37
7.29	5.69	4.95	4.64	4.72
7.86	6.16	5.36	5.02	5.09
8.44	6.65	5.80	5.44	5.51
9.02	7.12	6.20	5.79	5.81
9.57	7.63	6.63	6.18	6.18
10.1	8.17	7.13	6.64	6.56
10.7	8.58	7.49	6.99	6.95
11.2	9.06	7.94	7.42	7.38
11.8	9.55	8.38	7.83	7.78
12.3	10.2	8.84	8.27	8.20
12.8	10.5	9.28	8.68	8.60
13.4	11.0	9.73	9.11	9.01
13.9	11.5	10.2	9.53	9.43
14.4	12.0	10.6	9.96	9.85
	0.87 1.48 2.06 2.63 3.21 3.78 4.36 4.95 5.54 6.12 6.71 7.29 7.86 8.44 9.02 9.57 10.1 10.7 11.2 11.8 12.3 12.8 13.4 13.9	0.87       0.70         1.48       1.19         2.06       1.64         2.63       2.08         3.21       2.52         3.78       2.95         4.36       3.39         4.95       3.84         5.54       4.29         6.12       4.75         6.71       5.21         7.29       5.69         7.86       6.16         8.44       6.65         9.02       7.12         9.57       7.63         10.1       8.17         10.7       8.58         11.2       9.06         11.8       9.55         12.3       10.2         12.8       10.5         13.4       11.0         13.9       11.5	0.87       0.70       0.64         1.48       1.19       1.08         2.06       1.64       1.49         2.63       2.08       1.87         3.21       2.52       2.25         3.78       2.95       2.62         4.36       3.39       2.99         4.95       3.84       3.37         5.54       4.29       3.75         6.12       4.75       4.15         6.71       5.21       4.54         7.29       5.69       4.95         7.86       6.16       5.36         8.44       6.65       5.80         9.02       7.12       6.20         9.57       7.63       6.63         10.1       8.17       7.13         10.7       8.58       7.49         11.2       9.06       7.94         11.8       9.55       8.38         12.3       10.2       8.84         12.8       10.5       9.28         13.4       11.0       9.73         13.9       11.5       10.2	0.87         0.70         0.64         0.63           1.48         1.19         1.08         1.06           2.06         1.64         1.49         1.45           2.63         2.08         1.87         1.81           3.21         2.52         2.25         2.16           3.78         2.95         2.62         2.51           4.36         3.39         2.99         2.85           4.95         3.84         3.37         3.20           5.54         4.29         3.75         3.55           6.12         4.75         4.15         3.91           6.71         5.21         4.54         4.27           7.29         5.69         4.95         4.64           7.86         6.16         5.36         5.02           8.44         6.65         5.80         5.44           9.02         7.12         6.20         5.79           9.57         7.63         6.63         6.18           10.1         8.17         7.13         6.64           10.7         8.58         7.49         6.99           11.2         9.06         7.94         7.42

The symbol X denotes the mole fraction of  $^{3}\mathrm{P}$  atoms.

Table 35

Mole Fractions of the Atmospheric and Ablative Species During Jovian

Entry for Stagnation-Point Peak Heating

Species	Inner Boundary	Outer Boundary
Н	G.290	0.711
H <sup>+</sup>		0.118
Не	0.007	0.051
е	0.003	0.119
С	0.653	0.001
c <sub>2</sub>	0.010	
0	0.010	
CO	0.024	
c <sup>+</sup>	0.003	0.001
Total	1.000	1.001

 $\begin{tabular}{ll} Table & 36 \\ \hline \begin{tabular}{ll} Two Body Interactions at the Inner and Outer Mixing Boundaries \\ \hline \end{tabular}$ 

Inner Boundary	Outer Boundary
C-C	н-н
н-н	e-e
0-0	H <sup>+</sup> +H <sup>+</sup>
С-Н	Не-Не
C-0	Н-е
H-0	H-H <sup>+</sup>
	н-не
	e-H <sup>+</sup>
	е-Не
	H <sup>+</sup> -He

Table 37 Interaction Potential Parameters for the  $O(^3P)-O(^3P)$  Interaction

State	Method	<u>Potential</u>	Temperatures	Parameters
$3_{\Sigma_{\mathbf{q}}}$	RKR	AIP	a11	A=194.5 B=7.83
3 <sub>\(\Sigma\)</sub> g 1 \(\Delta\) g + g 1 \(\Sigma\) 1 \(\Sigm	RKR,HH	AIP	a11	A=123.4 B=7.89
$1_{\Sigma_{\mathbf{q}}^{+}}$	RKR,HH	AIP	all	A=141.6 B=8.64
1 <sub>Σ</sub> -	RKR,HH	ES	all	$\alpha$ =12 r <sub>e</sub> =1.597 $\epsilon$ =0.6655
$3\Sigma_{u}^{+}$	RKR	AIP	low	A=480.6 B=11.46
		ES	high	$\alpha$ =12 $r_e$ =1.518 $\epsilon$ =0.8239
$^{3}\Delta_{\mathbf{u}}$	RKR,HH	AIP	low	A=560.8 B=9.44
		ES	high	$\alpha$ =12 $r_e$ =1.480 $\epsilon$ =0.9157
3 <sub>11</sub> u	РР	ER	all	F=339 D=3.570
1 <sub>II</sub> u				
5 <sub>Π</sub> α	PP	ER	a11	F=717 D=3.565
<sup>5</sup> πg <sup>3</sup> πg <sup>1</sup> πg <sup>5</sup> Σ- <sup>1</sup> π2				
5 <sub>5</sub> -	PP	FP.	all	F=1057 D=3.567
<sup>2</sup> u 3 <sub>Σ</sub> + 2	• •	• •	411	1 1037 5 3.307
1 <sub>Σ</sub> + g2	PP	ER	all	F=1358 D=3.570
<sup>5</sup> Δg	PP	ER	all	F=1433 D=3.565
5 <sub>Σ</sub> + g				
<sup>5</sup> Σ <sup>+</sup> g  5Σ <sup>+</sup> σ  5Σ <sup>+</sup> σ  5π <sub>u</sub>	PP	ER	all	F=2114 D=3.567
5 <sub>11</sub> u	PP	ER	a11	F=2455 D=3.567

The parameters have been chosen so that V(r) is in electron volts and r is in Angstroms.

The symbol PP denotes the perfect pairing method.

Tab1n 38

Transport Collision	Integrals for the	$0(^{3}P)-0(^{3}P)$ Interaction
Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ ( $\Lambda^2$ )	$\sigma^2\Omega(2,2)$ * (Å <sup>2</sup> )
1	6.193	7.182
2	5.271	6.139
3	4.746	5.567
4	4.386	5.174
5	4.117	4.876
6	3.905	4.640
7	3.727	4.438
8	3.575	4.271
9	3.442	4.118
10	3.335	3.996
11	3.236	3.883
12	3.148	3.782
13	3.068	3.692
14	2.996	3.609
15	2.931	3.534
16	2.931	3.457
17	2.881	3.401
18	2.835	3.349
19	2.792	3.300
20	2.752	3.255
21	2.714	3.212
22	2.679	3.172
23	2.645	3.134
24	2.614	3.098
25	2.584	3.064

State	Potential	Parameters
с-н ( <sup>2</sup> п)	MP	C=1.547 $\epsilon$ =3.63 $r_e$ =1.120
C-H $(^4\Sigma)$	MP	(=1.913 $\epsilon$ =2.84 $r_e$ =1.086
C-H ( <sup>2</sup> Σ)	MP	C=4.080 $\epsilon$ =0.40 $r_e$ =1.164
с-н ( <sup>4</sup> п)	ER	F=194.53 D=0.3611
н-о ( <sup>2</sup> п)	MP	C=1.568 $\varepsilon$ =4.63 $r_e$ =0.9706
H-0 $(^4\Sigma)$	MP	C=1.331 $\varepsilon$ =4.22 $r_e$ =1.0121
H-0 ( <sup>4</sup> π)		
H-0 ( $^4\Sigma$ )		

The parameters are chosen so that V(r) is in electron volts and r is in Angstroms.

 $\label{thm:condition} Table~40$  Transport Collision Integrals for the C-H Interaction

$Tx10^{-3}$ (°K)	$\sigma^2_{\Omega}(1,1)$ * $(\mathring{A}^2)$	$\sigma^2\Omega^{(2,2)*}$ ( $\mathring{A}^2$ )	A*	B*
1	8.9178	8.8958	1.0476	1.1807
2	7.3553	7.3080	1.0543	1.2066
3	6.4580	6.4967	1.0672	1.2354
4	5.8014	5.9626	1.0840	1.2659
5	5.2802	5.5599	1.1023	1.2915
6	4.8347	5.2158	1.1198	1.3106
7	4.4532	4.9091	1.1357	1.3229
8	4.1314	4.6392	1.1497	1.3297
9	3.8412	4.3851	1.1623	1.3324
10	3.5931	4.1566	1.1728	1.3319
11	3.3715	3.9455	1.1819	1.3299
12	3.1821	3.7564	1.1891	1.3251
13	3.0094	3.5790	1.1953	1.3232
14	2.8527	3.4136	1.2003	1.3199
15	2.7140	3.2625	1.2041	1.3169
16	2.5917	3.1257	1.2068	1.3145
17	2.4820	3.0001	1.2086	1.3132
18	2.3786	2.8795	1.2096	1.3129
19	2.2932	2.7779	1.2100	1.3137
20	2.2062	2.6722	1.2093	1.3160
21	2.1363	2.5866	1.2086	1.3188
22	2.0645	2.4969	1.2070	1.3238
23	2.0066	2.4234	1.2052	1.3290
24	1.9463	2.3454	1.2022	1.3372
25	1.8945	2.2778	1.1997	1.3454

Table 41
Transport Collision Integrals for the H-O Interaction

	1141139010 001113	Ton Insagnats for t	one ii o iiice	action
$Tx10^{-3}$ (°K)	$\sigma^{2}\Omega^{(1,1)*}(\mathring{A}^{2})$	$\sigma^2\Omega^{(2,2)*}$ (Å <sup>2</sup> )	A*	B*
1	12.7841	11.0176	0.8636	1.1570
2	10.7138	9.1133	0.8521	1.1792
3	9.5289	8.1146	0.8525	1.2178
4	8.6459	7.4967	0.8674	1.2738
5	7.9063	7.0463	0.8911	1.3310
6	7.2735	6.6820	0.9183	1.3792
7	6.7162	6.3587	0.9462	1.4163
8	6.1758	6.0315	0.9759	1.4449
9	5.7258	5.7419	1.0019	1.4616
10	5.3050	5.4542	1.0272	1.4716
11	4.9335	5.1825	1.0496	1.4744
12	4.5980	4.9231	1.0698	1.4735
13	4.2928	4.6725	1.0877	1.4690
14	3.9961	4.4153	1.1042	1.4620
15	3.7675	4.2061	1.1158	1.4551
16	3.5385	3.9886	1.1267	1.4475
17	3.3392	3.7904	1.1346	1.4406
18	3.1485	3.5929	1.1408	1.4346
19	2.9966	3.4295	1.1442	1.4307
20	2.8513	3.2681	1.1461	1.4283
21	2.7128	3.1091	1.1461	1.4281
22	2.5815	2.9582	1.1440	1.4304
23	2.4816	2.8309	1.1409	1.4347
24	2.3787	2.7001	1.1354	1.4432
25	2.2888	2.5835	1.1292	1.4535

Table 42
Interaction Potential Parameters for the C-O Interaction

State	Potential	Para	meters	
1 <sub>2</sub> +	MP	C=1.902	€=11.242	r <sub>e</sub> =1.1281
1 <sub>Σ</sub> <sup>+</sup> <sub>2</sub> 3 <sub>Σ</sub> +				
5 <sub>Σ</sub> +				
5 <sub>Σ</sub> <sup>+</sup> <sub>2</sub>				
$1_{\Sigma}^{-1}$				
3 <sub>Σ</sub> -				
5 <sub>Σ</sub> -				
ιπ				
1 <sub>112</sub>				
3 <sub>11</sub>				
3 <sub>11</sub> 2				
5 <sub>II</sub>				
<sup>5</sup> 11 <sub>2</sub>				
10				
3				
5 <sub>Δ</sub> 3 <sub>Σ</sub> +				

The parameters are chosen so that V(r) is in electron volts and r is in Angstroms.

 $\label{thm:condition} Table~43$  Transport Collision Integrals for the C-O Interaction

2	0 (2 2)	יים און הוא האים המים	-u interact	ion
$T \times 10^{-3}$ (°K)	$\sigma^2\Omega^{(1,1)*}(\mathring{A}^2)$	$\sigma^{2}\Omega^{(2,2)*}$ ( $\mathring{A}^{2}$ )	A*	B*
1	14.6822	14.7130	1.0019	1.1922
2	12.8838	11.7392	0.9112	1.1183
3	11.8634	10.7609	0.9070	1.1579
4	11.1015	9.9934	0.9001	1.1711
5	10.5146	9.3813	0.8922	1.1742
6	10.0351	8.8927	0.8862	1.1779
7	9.6363	8.5095	0.8831	1.1851
8	9.2627	8.1775	0.8829	1.1970
9	8.9388	7.9137	0.8854	1.2121
10	8.6644	7.7073	0.8896	1.2283
11	8.3841	7.5108	0.8959	1.2477
12	8.1546	7.3590	0.9025	1.2654
13	7.9194	7.2101	0.9025	1.2846
14	7.6780	7.0621	0.9198	1.3052
15	7.4613	6.9318	0.9291	1.3239
16	7.2713	6.8183	0.9377	1.3402
17	7.0774	6.7023	0.9470	1.3565
18	6.8796	6.5831	0.9569	
19	6.7118	6.4805	0.9655	1.3726
20	6.5414	6.3747	0.9745	1.3857
21	6.3688	6.2652	0.9837	1.3983
22	6.1940	6.1518	0.9931	1.4103
23	6.0528	6.0581	1.0008	1.4216
24	5.8749	5.9372	1.0106	1.4299
25	5.7316	5.8372	1.0184	1.4395
		_ , , , , ,	1.0104	1.4463

Table 44

Mole Fractions of Species at the Surface of the Entry Probe

During Jovian Entry for Stagnation-Point Peak Heating

Species	Mole Fraction
c <sub>3</sub>	0.303
Н	0.149
C <sub>4</sub> H	0.138
C <sub>2</sub> H	0.115
со	0.108
с <sub>3</sub> н	0.099
$c_2$	0.028
С	0.026
H <sub>2</sub>	0.024
с <sub>2</sub> н <sub>2</sub>	0.010
Total	1.000

Table 45

Transport Collision Integrals for the CO-CO Interaction

Transport Coili	sion integrals for the	
$Tx10^{-3}$ (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$
1	7.9131	8.7386
2	7.1140	7.9304
3	6.6994	7.4950
4	6.4173	7.1974
5	6.2154	6.9724
6	6.0568	6.7999
7	5.9256	6.6571
8	5.8179	6.5386
9	5.7251	6.4348
10	5.6446	6.3461
11	5.5750	6.2703
12	5.5493	6.2429
13	5.5282	6.2214
14	5.5168	6.2136
15	5.5189	6.2141
16	5.5264	6.2262
17	5.5409	6.2458
18	5.5645	6.2720
19	5.5901	6.3033
20	5.6250	6.3442
21	5.6635	6.3889
22	5.7037	6.4404
23	5.7527	6.4933
24	5.8026	6.5522
25	5.8581	6.6167

 ${\it Table~46}$  Interaction Potential Energy for the He-C  $_{\rm 2}{\rm H~Interaction}$ 

r (Å)	V (He-C <sup>1</sup> ) <sub>av</sub>	V (He-C <sup>2</sup> ) <sub>av</sub>	V (He-H <sup>3</sup> ) <sub>av</sub>	V (He-C <sub>2</sub> H) <sub>av</sub>	V (ER)
1.85	3.18	0.30	0.59	4.07	3.88
1.90	2.32	0.25	0.44	3.01	3.01
1.95	1.73	0.22	33	2.28	2.33
2.00	1.31	0.19	0.25	1.75	1.81
2.05	1.01	0.16	0.20	1.36	1.40
2.10	0.78	0.14	0.15	1.07	1.09
2.15	0.62	0.12	0.12	0.86	0.84
2.18	0.54	0.11	0.11	0.75	0.72

ii diispoi c	Total store and agriculture of the the ogni	111001 4001011
$Tx10^{-3}$ (°K)	$\sigma^2\Omega^{(1,1)*}(\mathring{A}^2)$	$\sigma^2\Omega^{(2,2)*}(\mathring{A}^2)$
1	6.0301	6.8363
2	5.3911	6.1483
3	5.0422	5.7696
4	4.7966	5.5025
5	4.6144	5.3040
6	4.4599	5.1355
7	4.3399	5.0043
8	4.2371	4.8919
9	4.1433	4.7894
10	4.0584	4.6963
11	3.9895	4.6208
12	3.9212	4.5459
13	3.8610	4.4798
14	3.8087	4.4224
15	3.7568	4.3653
16	3.7053	4.3086
17	3.6614	4.2603
18	3.6250	4.2203
19	3.5832	4.1725
20	3.5455	4.1328
21	3.5097	4.0933
22	3.4812	4.0619
23	3.4458	4.0228
24	3.4175	3.9916
25	3.3894	3.9606

Table 48 The Binary Diffusion Coefficient, D (cm $^2$ /sec), for the He-C $_2$ H Interaction

Tx10 <sup>-3</sup> (°K)	D (this report)	D (Esch, et al. 72)
1	5.244	4.628
2	16.59	14.58
3	32.59	28.54
4	52.74	45.96
5	76.62	66.51
6	104.2	89.95
7	134.9	116.1
8	163.9	144.8
9	206.1	176.0
10	246.4	209.6
11	289.2	245.4
12	335.2	283.5
13	383.9	323.6
14	434.9	365.9
15	489.0	410.1
16	546.2	456.4
17	605.4	535.4
18	666.2	554.8
19	730.9	606.7
20	797.8	660.5
21	867.1	716.1
22	937.4	773.4
23	1012	832.5
24	1088	893.3
25	1166	955.8

Table 49  $\label{eq:table 49} Interaction Potential Energy for the C-C_2 Interaction for the <math display="inline">^1\Sigma_q^+$  State of C\_2

	for the 2g state of c2	
r (Å)	v (r-c <sub>2</sub> ) <sub>av</sub>	V (Morse)
0.010	-874585	
0.200	-9174	
0.500	180.7	
0.550	231.5	
0.600	234.6	449.9
0.800	123.5	149.7
1.000	39.58	41.03
1.242	1.42	1.15
1.323	-3.10	-3.28
1.429	-6.02	-6.07
1.560	-6.93	-6.93
1.588	-6.90	-6.90
1.852	-5.17	-5.27
2.117	-3.14	-3.31
2.646	-0.96	-1.09
3.704	-0.07	-0.10
5.292	0	0
6.000	0	0

for the  $^1\Sigma_q^+$  State of  $\mathrm{C}_2$  $Tx10^{-3}$  (°K) В\* Α 1 13.9192 13.2872 0.9546 1.1074 2 12.1639 11.3227 0.9308 1.1541 3 10.2096 11.1142 0.9186 1.1660 4 10.3457 9.4087 0.9094 1.1715 5 9.7603 8.8495 0.9067 1.1846 6 9.2741 8.4349 0.9095 1.2051 8.8299 7 8.0963 0.9169 1.2317 8 8.4562 7.8359 0.9267 1.2587 8.0981 9 7.6006 7.9386 1.2871 10 7.7569 7.3835 0.9519 1.3151 11 7.4680 7.2009 0.9642 1.3386 12 7.1701 7.0108 0.9778 1.3617 13 6.8982 6.8335 0.9906 1.3813 14 6.6705 1.0023 1.3974 6.6551 15 6.4077 6.4996 1.0143 1.4119 6.3201 1.0265 1.4246 16 6.1568 1.4339 17 5.9395 6.1593 1.0370 18 6.0204 5.7574 1.0456 1.4402 19 5.5380 5.8481 1.0559 1.4462 20 5.3551 5.6999 1.0644 1.4497 21 5.5480 1.0726 1.4517 5.1725 5.4240 1.0790 1.45^ 22. 5.0269 23 4.8459 5.2663 1.0867 1.4520 1.0927 1.4506 24 4.7023 5.1382 25 4.5598 5.0088 1.0984 1.4484

Table 51 Interaction Potential Energy for the  ${\rm C_2-C_2}$  Interaction for the  $^1{\rm E}_g^+$  State of  ${\rm C_2}$ 

_	y	2
r (Å)	$v(c_2-c_2)_{av}$	V(Morse)
0.100	-732400	
0.300	-65574	
0.600	-3393	
0.900	10.05	
1.000	81.04	
1.058	83.55	188.3
1.323	27.97	34.99
1.429	11.41	13.19
1.588	-2.21	-2.16
1.852	<b>-</b> 7.92	-7.91
1.867	<del>-</del> 7.92	-7.92
2.117	-6.41	-6.49
2.646	-2.34	-2.58
3.704	-0.19	-0.27
4.000	-0.09	-0.14
4.292	0	-0.07
5.000	0.01	-0.02
6.000	0	0

Table 52

Transport Collision Integrals for the  $C_2$ - $C_2$  Interaction for the  $l_x$ + state of c

	the $^{1}\Sigma_{g}^{+}$ State of $^{0}\Sigma_{g}$	
$T \times 10^{-3}$ (°K)	$\sigma^2\Omega^{(1,1)*}(\mathring{A}^2)$	$\sigma^2\Omega(2,2)*(\mathring{A}^2)$
1	18.6803	20.5068
2	16.4601	17.0317
3	15.2878	15.5781
4	14.4943	14.7858
5	13.8734	14.1219
i	13.3545	13.5339
7	12.8881	13.0009
8	12.5060	12.5760
9	12.1494	12.1983
10	11.8476	11.8963
11	11.5397	11.6059
12	11.2875	11.3806
13	11.0290	11.1597
14	10.7630	10.9413
15	10.5234	10.7502
16	10.3126	10.5851
17	10.0966	10.4176
18	9.8753	10.2466
19	9.6867	10.1008
20	9.4945	9.9513
21	9.2988	9.7977
22	9.0997	9.6397
23	8.9382	9.5098
24	8.7338	9.3431
25	8.5634	9.2060

Table 53  $\label{eq:comparison} \text{Comparison of the C-C, C-$\mathbb{C}_2$, and $\mathbb{C}_2$-$\mathbb{C}_2$ Interaction } \\ \text{Potentials Corresponding to $^+$he $^1\Sigma_g^+$ State of $\mathbb{C}_2$}$ 

			, .,
r(Å)	V(C-C)	v(c-c <sub>2</sub> )	v(c <sub>2</sub> -c <sub>2</sub> )
1.058	-4.42	25.68	83.55
1.111	-5.49	16.05	75.75
1.242	-6.36	1.42	45.01
1.323	-6.15	-3.10	27.97
1.429	-5.47	-6.02	11.41
1.588	-4.15	~6.90	-2.21
1.852	-2.26	-5.17	-7.92
2.117	-1.13	-3.14	-6.41
2.646	-0.29	-0.96	-2.34
3.704	-0.03	-0.07	-0.19
5.292	0	0	0

Table 54 The Translational Contribution to the Thermal Conductivity,  $\lambda_{\rm tr}^{\rm mix}(\rm W/m/^{\circ}K)$  , for a Mixture of C and C  $_2$ 

$Tx10^{-3}(°K)$	X=0.00	X=0.25	X=0.50	X=0.75	X=1.00
4	003	0.117	0.134	0.157	0.195
5	0.120	0.139	0.159	0.187	0.231
6	0.137	0.159	0.184	0.216	0.267
7	0.155	0.180	0.207	0.244	0.301
8	0.177	0.199	0.230	0.271	0.334

The symbol X denotes the mole fraction of carbon atoms.

 $$\mathsf{Table}$$  55 Transport Collision Integrals for the He-O Interaction

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2.2)*}(\mathring{A}^2)$	A*	B*
1	2.3220	2.7032	1.1654	1.1279
2	2.0096	2.3584	1.1748	1.1368
3	1.8348	2.1645	1.1809	1.1427
4	1.7162	2.0322	1.1854	1.1471
5	1.6288	1.9346	1.1890	1.1507
6	1.5590	1.8564	1.1921	1.1538
7	1.4982	1.7883	1.1948	1.1566
8	1.4498	1.7339	1.1972	1.1589
9	1.4058	1.6843	1.1994	1.1612
10	1.3696	1.6436	1.2013	1.1631
11	1.3340	1.6034	1.2032	1.1651
12	1.3023	1.5675	1.2049	1.1669
13	1.2745	1.5360	1.2065	1.1686
14	1.2504	1.5087	1.2079	1.1700
15	1.2265	1.4817	1.2093	1.1715
16	1.2062	1.4586	1.2106	1.1728
17	1.1861	1.4358	1.2118	1.1741
18	1.1662	1.4131	1.2130	1.1754
19	1.1476	1.3907	1.2143	1.1768
20	1.1301	1.3720	1.2154	1.1780
21	1.1138	1.3536	1.2165	1.1791
22	1.1010	1.3389	1.2174	1.1801
23	1.0850	1.3206	1.2184	1.1812
24	1.0723	1.3061	1.2193	1.1822
25	1.0596	1.2916	1.2202	1.1832

Table 56 Transport Collision Integrals for the He-C Interaction

Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}$ (Å <sup>2</sup> )	$\sigma^2\Omega^{(2,2)*}$ ( $^2$ )
1	4.2921	5.2846
2	3.4359	4.2304
3	3.0166	3.7141
4	2.7505	3.3865
5	2.5604	3.1524
6	2.4148	2.9732
7	2.2982	2.8297
8	2.2018	2.7110
9	2.1201	2.6104
10	2.0496	2.5236
11	1.9879	2.4475
12	1.9331	2.3801
13	1.8841	2.3197
14	1.8398	2.2652
15	1.7995	2.2156
16	1.7626	2.1702
17	1.7286	2.1283
18	1.6972	2.0896
19	1.6680	2.0537
20	1.6407	2.0202
21	1.6152	1.9888
22	1.5913	1.9593
23	1.5688	1.9315
24	1.5475	1.9053
25	1.5273	1.8805

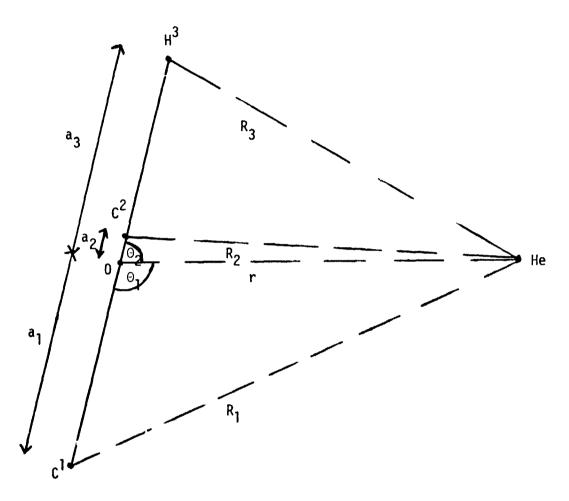
 $A^* = 1.2312$   $B^* = 1.1792$ 

Transport correspond integrals for the co-in interaction						
Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega^{(1,1)*}(\mathring{A}^2)$	$\sigma^2\Omega^{(2,2)*}$ (Å <sup>2</sup> )	A*	<u>B</u> *		
1	11.0850	10.2919	0.9675	1.1640		
2	9.0892	8.4479	0.9759	1.2266		
3	7.9100	7.5014	1.0004	1.2822		
4	6.9759	6.8435	1.0295	1.3278		
5	6.2486	6.3261	1.0571	1.3585		
6	5.6166	5.8630	1.0832	1.3776		
7	5.0940	5.4561	1.1052	1.3860		
8	4.6573	5.0982	1.1238	1.3881		
9	4.2643	4.7581	1.1401	1.3854		
10	3.9273	4.4508	1.1535	1.3804		
11	3.6442	4.1795	1.1639	1.3748		
12	3.3872	3.9230	1.1722	1.3688		
13	3.1676	3.6944	1.1783	1.3639		
14	2.9826	3.4945	1.1821	1.3606		
15	2.8099	3.3024	1.1844	1.3586		
16	2.6554	3.1244	1.1840	1.3587		
17	2.5238	2.9688	1.1838	1.3609		
18	2.4124	2.8341	1.1820	1.3648		
19	2.2977	2.6813	1.1781	1.3720		
20	2.2033	2.5707	1.1738	1.3813		
21	2.1191	2.4604	1.1680	1.3932		
22	2.0511	2.3696	1.1626	1.4059		
23	1.9742	2.2639	1 1546	1.4250		
24	1.9163	2.1824	1.1473	1.4440		
25	1.8616	2.1038	1.1391	1.4666		

i i alispoi c	corrision incegrars for	the C-C Inte	ELACTION
Tx10 <sup>-3</sup> (°K)	$\sigma^2\Omega(2,2)$ * ( $\mathring{A}^2$ )	A*	
1	6.9224	1.1257	1.1017
2	6.1607	1.1118	1.1012
3	5.7784	1.1049	1.1178
4	5.4646	1.0971	1.1199
5	5.2200	1.0896	1.1182
6	5.0279	1.0851	1.1214
7	4.8674	1.0824	1.1310
8	4.7462	1.0817	1.1433
9	4.6358	1.0828	1.1570
10	4.5398	1.0840	1.1709
11	4.4492	1.0861	1.1849
12	4.3667	1.0888	1.1976
13	4.2884	1.0914	1.2091
14	4.2142	1.0942	1.2193
15	4.1408	1.0971	1.2282
16	4.0728	1.0998	1.2356
17	4.0030	1.1024	1.2423
18	3.9364	1.1050	1.2476
19	3.8705	1.1074	1.2519
20	3.8076	1.1095	1.2552
21	3.7458	1.1114	1.2574
22	3.6875	1.1131	1.2588
23	3.6268	1.1146	1.2598
24	3.5706	1.1160	1.2598
25	3.5161	1.1172	1.2599

Figure 1

## Coordinate System for the $\mathrm{He-C_2H}$ Interaction



0 = center of geometry

$$a_1 = a_3 = 1.134 \text{ Å}$$
  $a_2 = 0.073 \text{ Å}$   $\Theta_2 = \Theta_3$ 

## <u>Publications</u>

- L. Biolsi
  "Transport Properties in the Jovian Atmosphere"
  Journal of Geophysical Research, <u>83</u>, 1125 (1978).
- L. Biolsi
  "Transport Properties of Monatomic Carbon"
  Journal of Geophysical Research, 83, 2476 (1978).
- L. Biolsi and L. R. Wallace "Some Effects of Ablation on Transport Properties in the Jovian Atmosphere" Progress in Astronautics and Aeronautics: Thermal Protection Systems (in press).

## Presentations

- L. Biolsi
  "Effect of Surface Ablation on Transport Properties in a H<sub>2</sub>-He Atmosphere'
  173rd National Meeting, American Chemical Society
  New Orleans, LA, March 20-25, 1977.
- L. Biolsi "Transport Properties in the Atmosphere of Jupiter" 13th Midwest Regional Meeting, American Chemical Society Rolla, MO, Nov. 3-4, 1977.
- L. Biolsi "Transport Properties of Gaseous Carbon" 33rd Southwest Regional Meeting, American Chemical Society Little Rock, AR, Dec. 5-7, 1977.
- L. Biolsi
  "Some Efrects of Ablation on Transport Properties in the Jovian Atmosphere"
  2nd AIAA/ASME Thermophysics and Heat Transfer Conference
  Palo Alto, CA, May 24-26, 1978.
- L. Biolsi and K. J. Biolsi "The Thermal Conductivity of Carbon in Electronically Excited States" 14th Midwest Regional Meeting, American Chemical Society Fayetteville, AR, Oct. 26-27, 1978
- L. Biolsi and L. R. Wallace "Transport Properties in a C<sub>2</sub>-C System" 14th Midwest Regional Meeting, American Chemical Society Fayetteville, AR, Oct. 26-27, 1978.